

A NuFact Study including all Parameter Correlations

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hep-ph/0105071 M. Freund, P. Huber and M. Lindner

- Analytical formulas for oscillation in matter
- Leading Parameters: Δm_{31}^2 and θ_{23}
- Sub-Leading Parameters: θ_{13} and $\text{sgn}\Delta m_{31}^2$
- Small- Δm^2 Effects: Δm_{21}^2 , θ_{21} and δ_{CP}
- Conclusions

Mixing Parameters in Matter

Expansion in the mass hierarchy $\alpha = \Delta m_{21}^2 / \Delta m_{31}^2$
and assuming small θ_{13} :

(M. Freund, hep-ph/0103300.)

$$\sin^2 2\theta'_{13} = \frac{\sin^2 2\theta_{13}}{\hat{C}^2} + \alpha \frac{2\hat{A}(-\hat{A} + \cos 2\theta_{13}) \sin^2 \theta_{12} \sin^2 2\theta_{13}}{\hat{C}^4}$$

$$\sin 2\theta'_{12} = \alpha \frac{2\hat{C} \sin 2\theta_{12}}{|\hat{A}| \cos \theta_{13} \sqrt{2\hat{C}(\mp \hat{A} + \hat{C} \pm \cos 2\theta_{13})}}$$

$$\sin 2\theta'_{23} = \sin 2\theta_{23} + \alpha \cos \delta \frac{2\hat{A} \sin 2\theta_{12} \sin \theta_{13} \cos 2\theta_{23}}{\pm 1 + \hat{C} \mp \hat{A} \cos 2\theta_{13}}$$

$$\sin \delta' = \sin \delta \left(1 - \alpha \frac{\cos \delta}{\tan 2\theta_{23}} \frac{2\hat{A} \sin 2\theta_{12} \sin \theta_{13}}{\pm 1 + \hat{C} \mp \hat{A} \cos 2\theta_{13}} \right)$$

with $\hat{C} = \sqrt{(\hat{A} - \cos 2\theta_{13})^2 + \sin^2 2\theta_{13}}$

$$\hat{A} = A / \Delta m_{31}^2 = 2VE / \Delta m_{31}^2$$

$$V = \sqrt{2}G_F n_e.$$

$$\Rightarrow \begin{aligned} \sin^2 2\theta'_{13} &\approx \frac{\sin^2 2\theta_{13}}{\hat{C}^2} \\ \sin^2 2\theta'_{23} &\approx \sin^2 2\theta_{23} \\ \sin \delta' &\approx \sin \delta \\ J'_{\text{CP}} &= \frac{\alpha}{|\hat{A}| \hat{C} \cos^2 \theta_{13}} J_{\text{CP}} \end{aligned}$$

3ν-Formulas in Matter

Probabilities in matter expanded in the mass hierarchy
 $\alpha = \Delta m_{21}^2 / \Delta m_{31}^2$ and assuming small θ_{13} :

$$P(\nu_\mu \rightarrow \nu_\mu) \approx$$

$$1 - \cos^2 \theta_{13} \sin^2 2\theta_{23} \sin^2 \Delta$$

$$+ 2\alpha \cos^2 \theta_{13} \cos^2 \theta_{12} \sin^2 2\theta_{23} \Delta \cos \Delta$$

$$P(\nu_e \rightarrow \nu_\mu) \approx$$

$$\sin^2 \theta_{23} \frac{\sin^2 2\theta_{13}}{(\hat{A} - 1)^2} \sin^2((\hat{A} - 1)\Delta)$$

$$\pm \alpha \frac{\sin \delta_{\text{CP}} \cos \theta_{13} \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23}}{\hat{A}(1 - \hat{A})} \sin(\Delta) \sin(\hat{A}\Delta) \sin((1 - \hat{A})\Delta)$$

$$+ \alpha \frac{\cos \delta_{\text{CP}} \cos \theta_{13} \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23}}{\hat{A}(1 - \hat{A})} \cos(\Delta) \sin(\hat{A}\Delta) \sin((1 - \hat{A})\Delta)$$

$$+ \alpha^2 \frac{\cos^2 \theta_{23} \sin^2 2\theta_{12}}{\hat{A}^2} \sin^2(\hat{A}\Delta)$$

$$\Delta = \Delta m_{31}^2 L / 4E \quad \hat{A} = A / \Delta m_{31}^2 = 2VE / \Delta m_{31}^2 \quad V = \sqrt{2}G_F n_e.$$

M. Freund, hep-ph/0103300.

A. Cervera *et al.*, Nucl. Phys. **B579** (2000) 17.

Framework

Simulation of event rate spectra

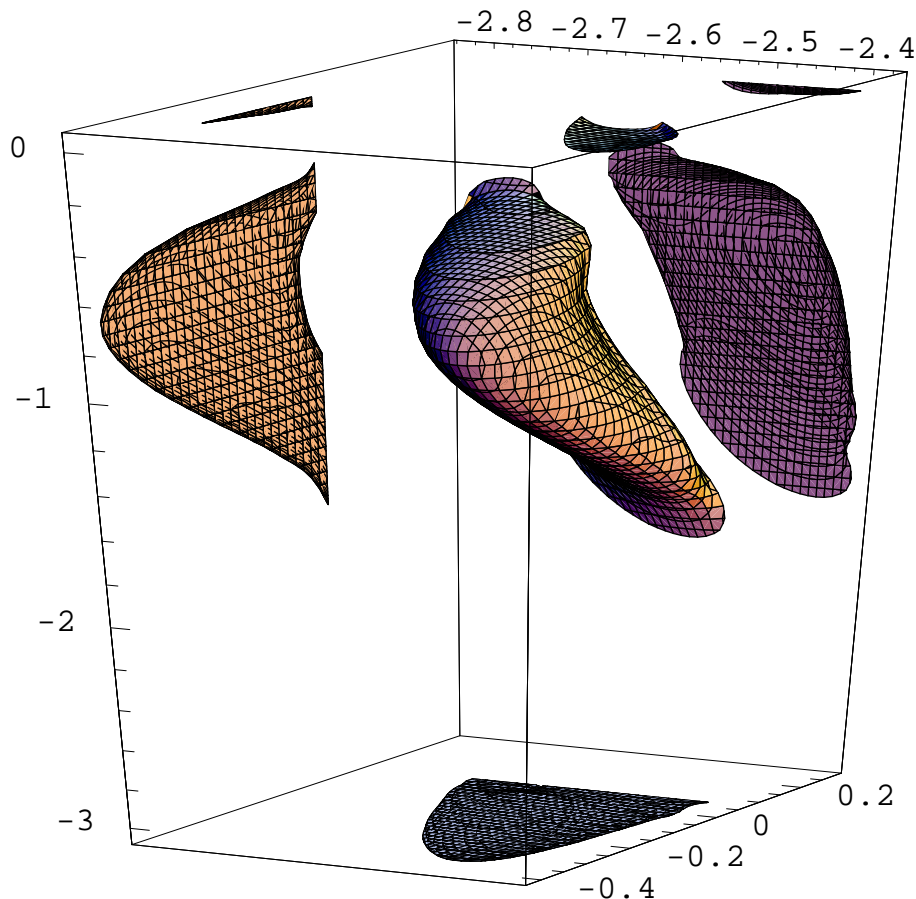
- **channels:** ν_μ -appearance and disappearance
- **flux:** $2 \cdot 10^{20}$ useful muon decays
- **detector:** 10 kt mass
perfect charge identification
Gaussian energy resolution ($\sigma = 0.1E$)
- **systematical errors and backgrounds:** not included

Physical input

- full numerical 3- ν calculation
- matter density profile taken into account
- central values for the simulation: $\theta_{23} = \theta_{12} = \pi/4$
 $\Delta m_{31}^2 = 3.5 \cdot 10^{-3} \text{ eV}^2$

Statistics

full six-parameter fits are too CPU-time consuming for scans of the parameter space!

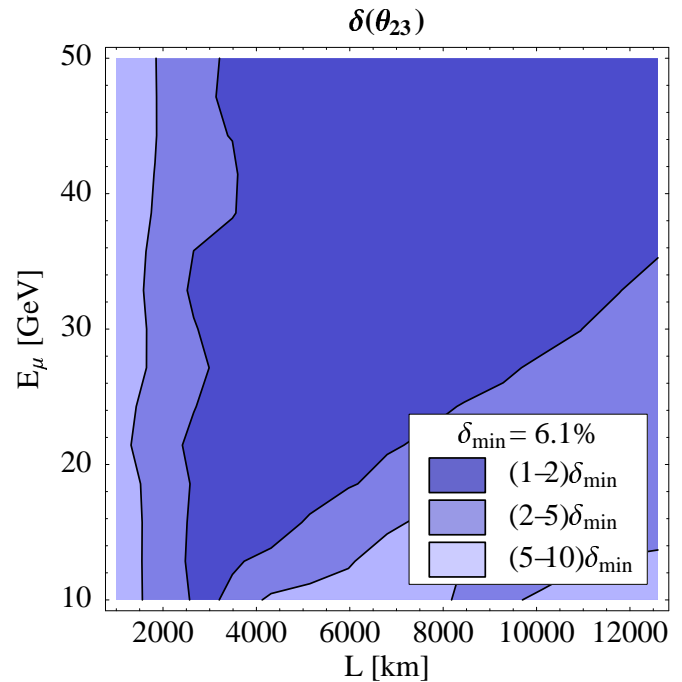
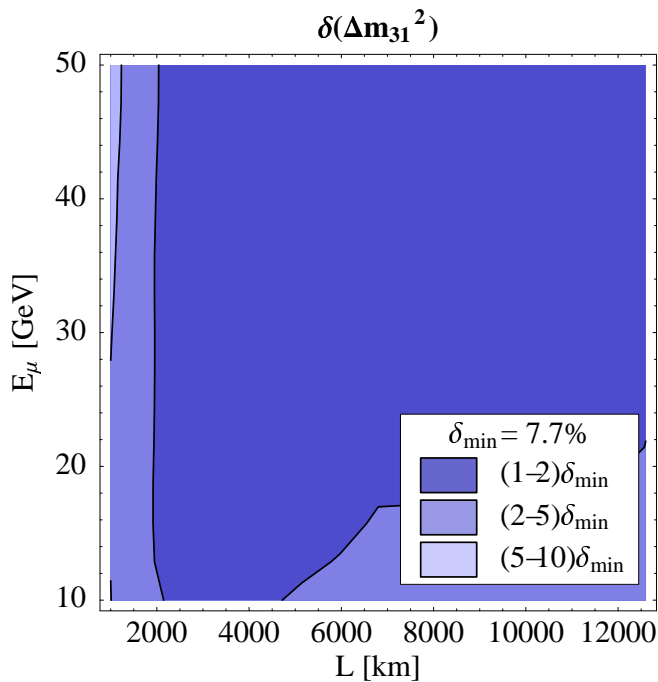


⇒ Calculation of all possible two-parameter fits

Advantage: automatically takes into account all **two parameter correlations**

Leading Parameters – $\Delta m_{31}^2, \theta_{23}$

L - E optimization



- Baseline should exceed a minimal value:

$$L > 3000 \text{ for } \Delta m_{31}^2 = 3.5 \cdot 10^{-3} \text{ eV}^2$$

$$L > 1000 \text{ for } \Delta m_{31}^2 = 5.0 \cdot 10^{-3} \text{ eV}^2$$

$$L > 5000 \text{ for } \Delta m_{31}^2 = 1.0 \cdot 10^{-3} \text{ eV}^2$$

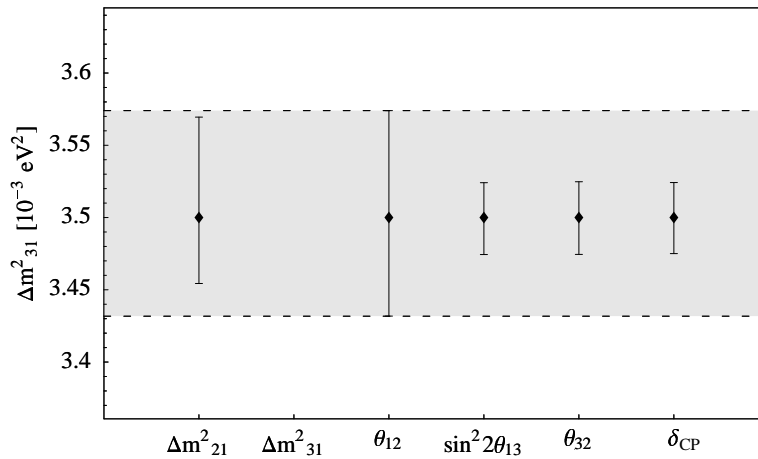
- $E_{\mu} > 30 \text{ GeV}$ preferable

- achievable precision for Δm_{31}^2 : $\delta_{\min} = 8\%$

- achievable precision for θ_{23} : $\delta_{\min} = 6\%$

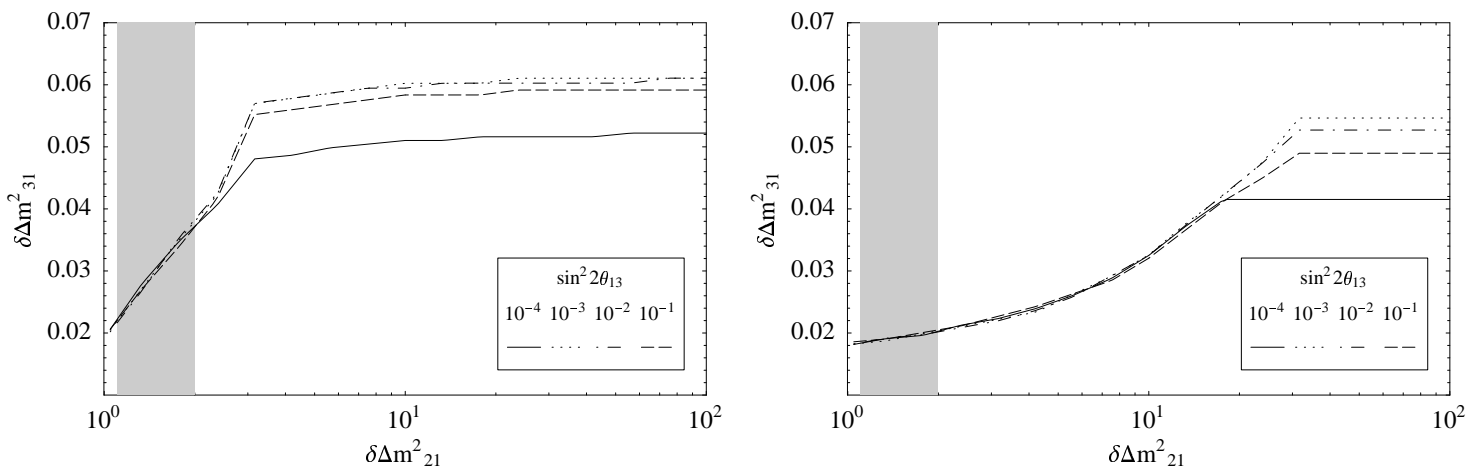
Leading Parameters – Δm_{31}^2

correlation with Δm_{21}^2



$E = 50 \text{ GeV}$, $L = 8000 \text{ km}$, $\sin^2 2\theta_{13} = 10^{-3}$, $\Delta m_{21}^2 = 10^{-4} \text{ eV}^2$

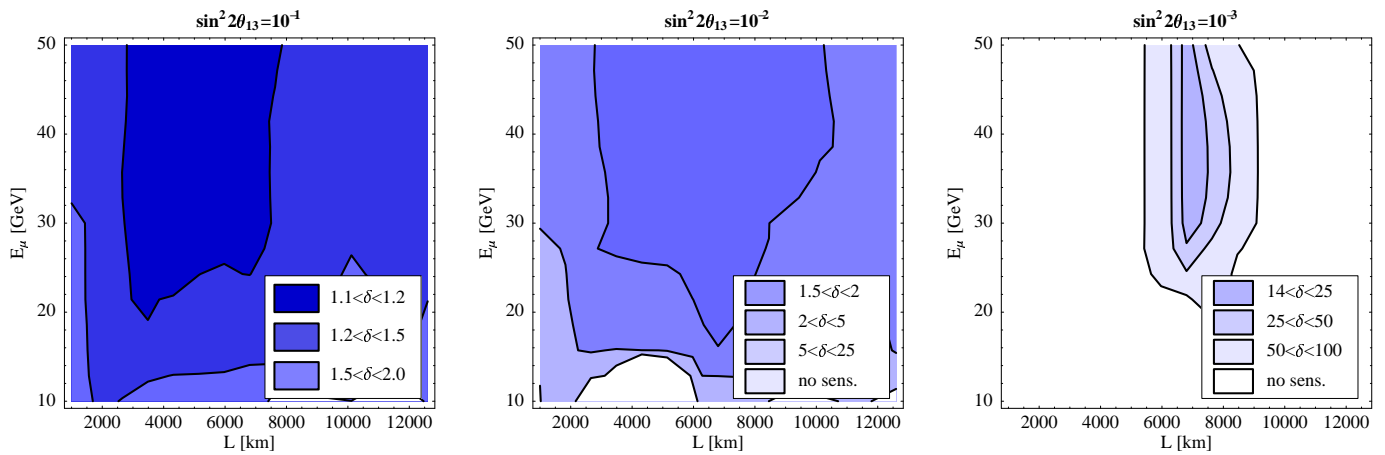
KamLand input



$\Delta m_{21}^2 = 10^{-4} \text{ eV}^2$ (left plot), $\Delta m_{21}^2 = 10^{-5} \text{ eV}^2$ (right plot)

Sub-Leading Parameters – $\sin^2 2\theta_{13}$

L - E optimization



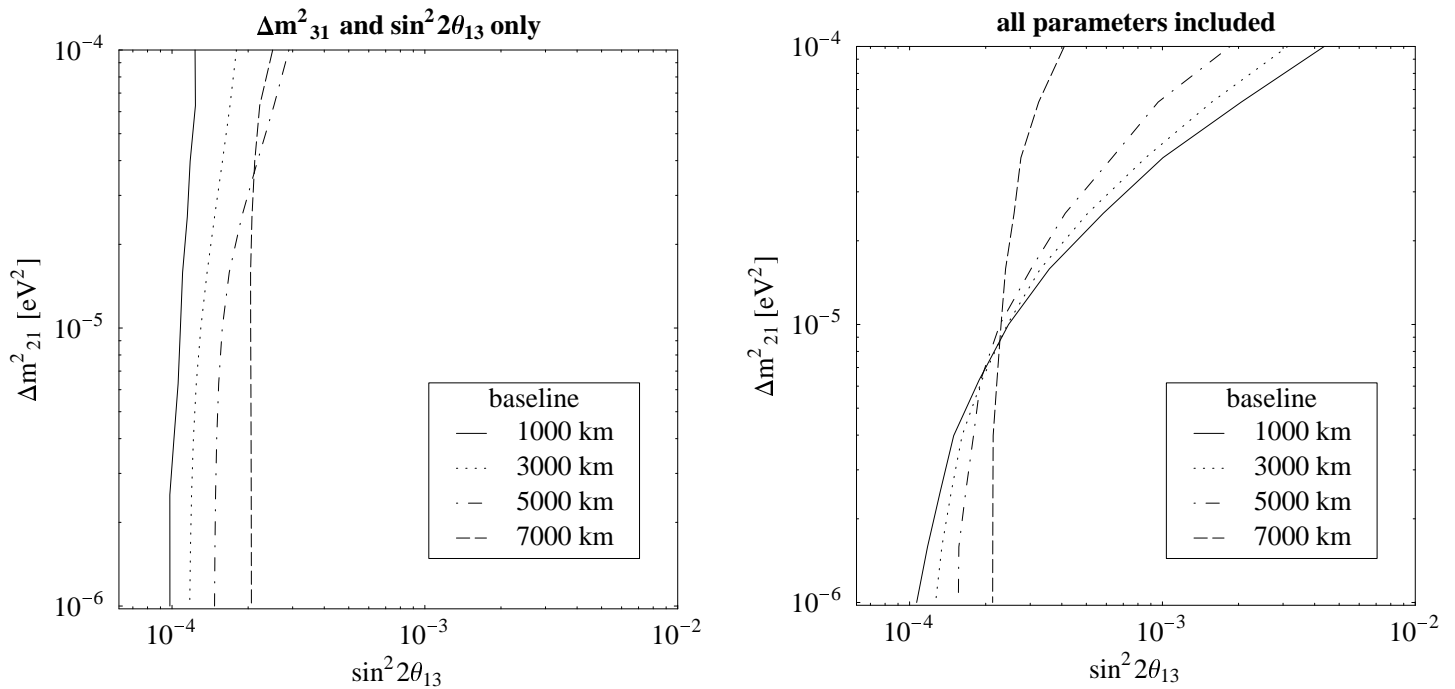
δ_{CP} unknown

- large θ_{13} : $3000 \text{ km} < L < 8000 \text{ km}$ similar
- $\sin^2 2\theta_{13} \lesssim 10^{-3}$: $7000 \text{ km} < L < 8000 \text{ km}$ best

small $\theta_{13} \Rightarrow$ large baseline

Sub-Leading Parameters – $\sin^2 2\theta_{13}$

$\sin^2 2\theta_{13}$ sensitivity limit



$$E_\mu = 50 \text{ GeV}$$

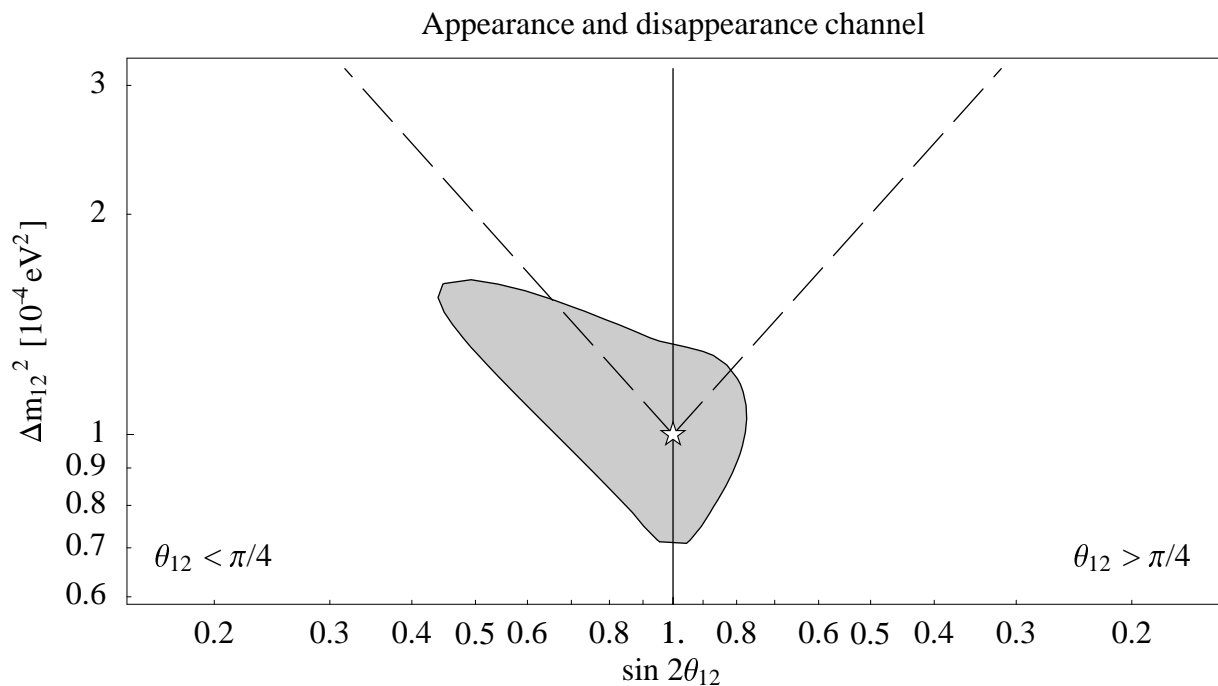
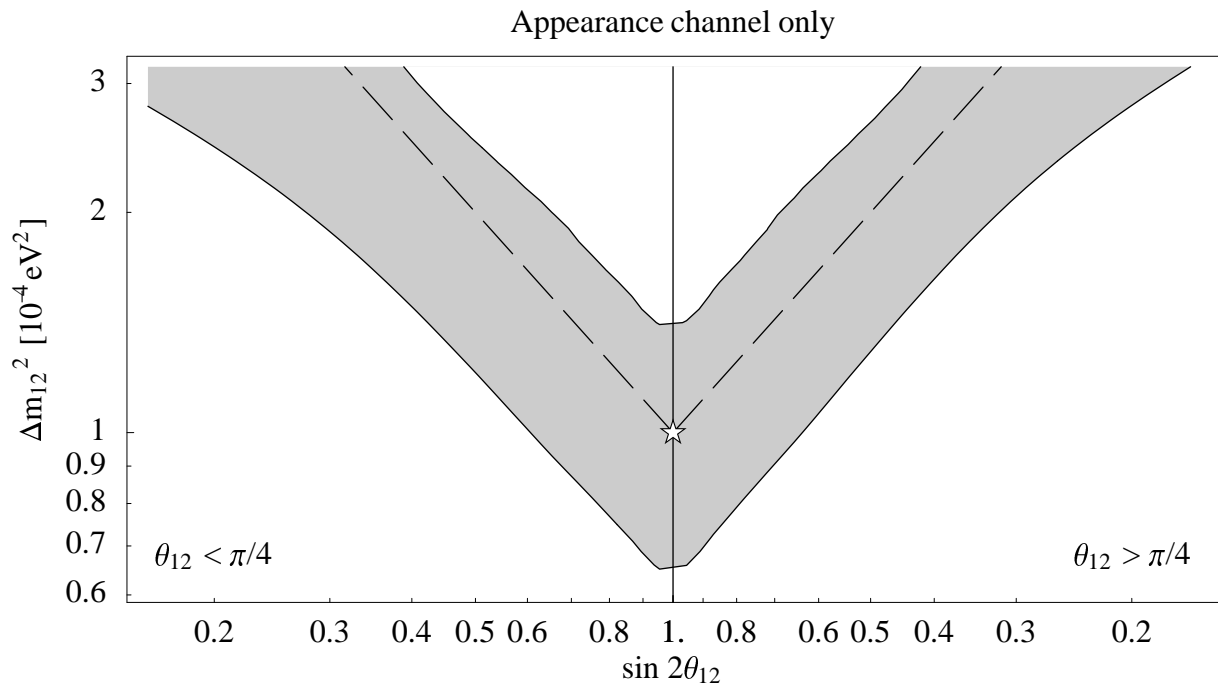
- left plot: only two parameter fit of $\sin^2 2\theta_{13}$ and Δm^2_{21}
- right plot: all correlations included

In the LMA-case, correlations with θ_{12} and δ_{CP} deteriorate sensitivity at high values of Δm^2_{21} .

(This effect is weaker at longer baselines.)

Small- Δm^2 effects – $\Delta m_{21}^2, \theta_{12}$

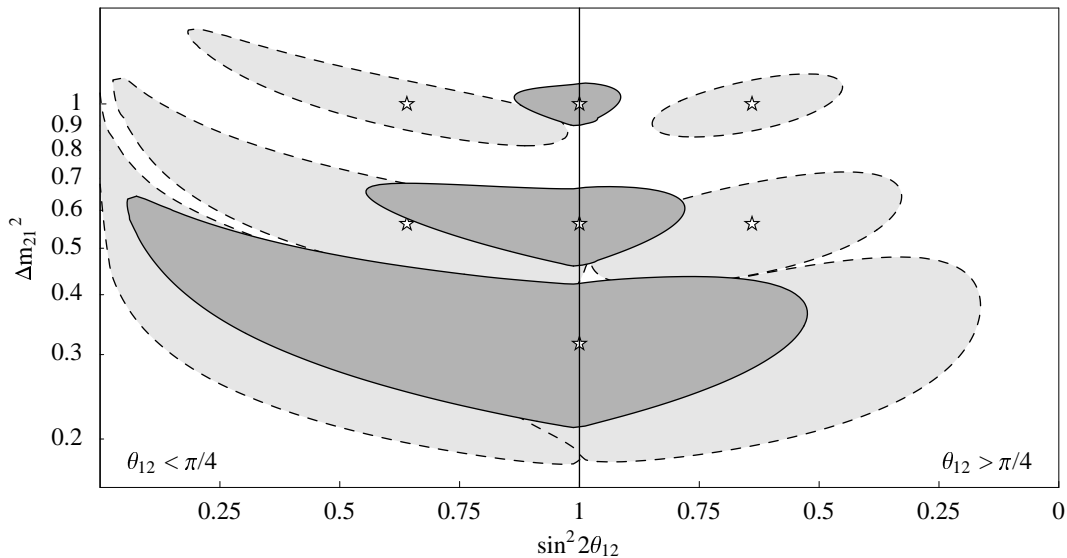
correlation of Δm_{21}^2 and θ_{12}



$E = 50 \text{ GeV}, L = 3000 \text{ km}$

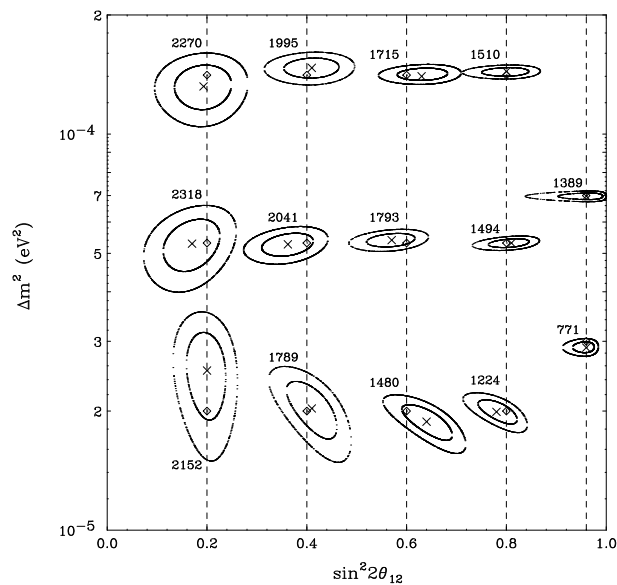
Small- Δm^2 effects – $\Delta m_{21}^2, \theta_{12}$

determination of Δm_{21}^2 and θ_{12}



$$E = 50 \text{ GeV}, L = 3000 \text{ km}, N_{\mu} m_{\text{kt}} = 2 \cdot 10^{22} (!)$$

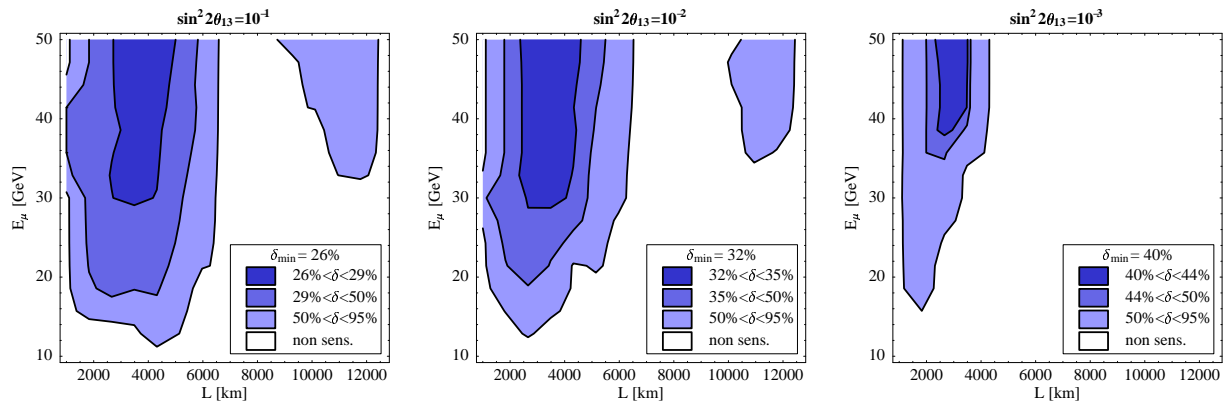
accuracy of KamLand:



V. Barger, D. Marfatia and B.P. Wood, Phys. Lett. **B498** (2001) 53.

Small- Δm^2 effects – δ_{CP}

L - E optimization for δ_{CP}

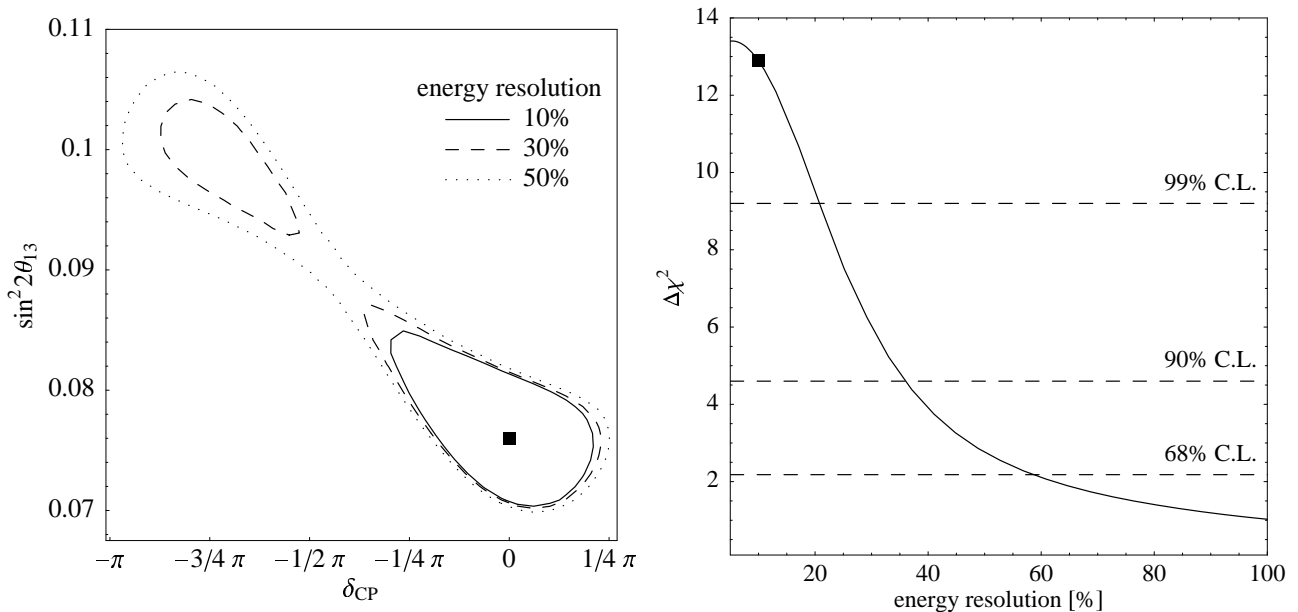


- baselines around 3000 km best
- higher energies preferred
- optimal baseline depends on the value of Δm_{31}^2

| Δm_{31}^2 | $\sin^2 2\theta_{13}$ | baseline | beam energy |
|----------------------------------|-----------------------|---|--------------------------------|
| $6.0 \cdot 10^{-3} \text{ eV}^2$ | 10^{-1} | $2400 \text{ km} \lesssim L \lesssim 4200 \text{ km}$ | $E_\mu \gtrsim 40 \text{ GeV}$ |
| | 10^{-2} | $2400 \text{ km} \lesssim L \lesssim 4000 \text{ km}$ | $E_\mu \gtrsim 40 \text{ GeV}$ |
| | 10^{-3} | $1800 \text{ km} \lesssim L \lesssim 2700 \text{ km}$ | $E_\mu \gtrsim 45 \text{ GeV}$ |
| $3.5 \cdot 10^{-3} \text{ eV}^2$ | 10^{-1} | $2800 \text{ km} \lesssim L \lesssim 4500 \text{ km}$ | $E_\mu \gtrsim 30 \text{ GeV}$ |
| | 10^{-2} | $2500 \text{ km} \lesssim L \lesssim 4500 \text{ km}$ | $E_\mu \gtrsim 30 \text{ GeV}$ |
| | 10^{-3} | $2500 \text{ km} \lesssim L \lesssim 3500 \text{ km}$ | $E_\mu \gtrsim 40 \text{ GeV}$ |
| $1.0 \cdot 10^{-3} \text{ eV}^2$ | 10^{-1} | no sens. | no sens. |
| | 10^{-2} | no sens. | no sens. |
| | 10^{-3} | no sens. | no sens. |

Small- Δm^2 effects – δ_{CP}

correlation of δ_{CP} with θ_{13}



$$E = 50 \text{ GeV}, L = 3000 \text{ km}, \Delta m_{21}^2 = 10^{-4} \text{ eV}^2$$

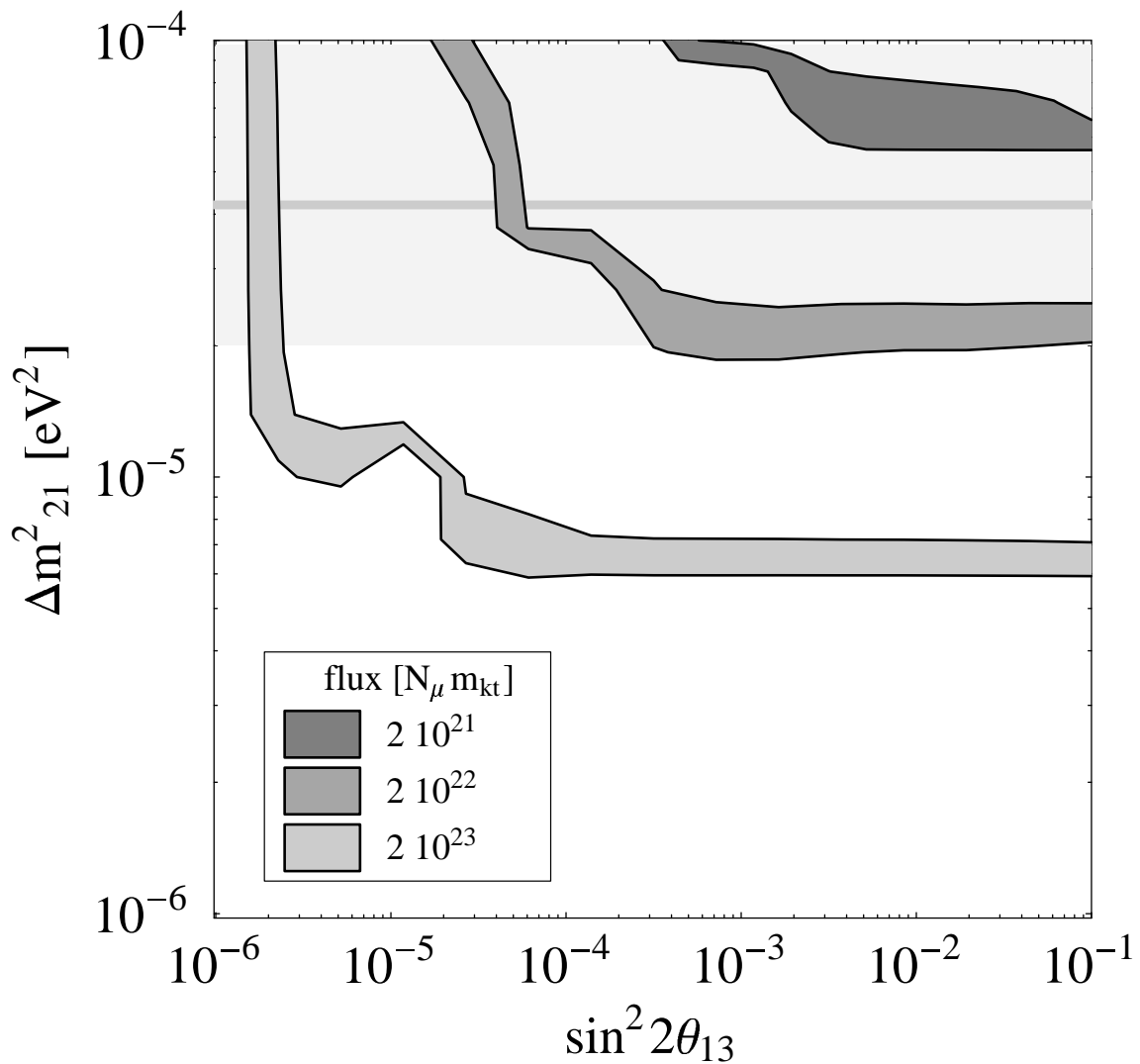
- two degenerate solutions¹
- depends strongly on energy resolution
- degeneracy lifted for energy resolution better than 25%

degeneracy \longleftrightarrow spectral information

¹J. Burguet-Castell *et al.*, hep-ph/0103258.

Small- Δm^2 effects – δ_{CP}

sensitivity limit for δ_{CP}



$E = 50 \text{ GeV}, L = 3000 \text{ km}$

$2 \cdot 10^{20} \text{ muons} \otimes 10 \text{ kt}$ is the minimum

Conclusions

Statistics

- the information in the event rate spectra should be optimally exploited
- parameter correlations are important

Leading Parameters

- errors on Δm_{31}^2 and θ_{23} about 8%
- $L_{min} \simeq 3\,000$ km
- $E > 30$ GeV

Sub-Leading Parameters

- strong correlation with Δm_{21}^2 or δ_{CP} depending on δ_{CP}
- higher baselines around 7 000 km are most sensitive
- KamLand input only for $\cos \delta_{CP} \geq 0$ helpful

Small- Δm^2 effects

- sensitivity to Δm_{21}^2 and θ_{12} is much worse than at e.g. KamLand
- sensitivity to δ_{CP} best at 50 GeV and $\simeq 3\,000$ km
- sufficient energy resolution lifts degeneracy between δ_{CP} and θ_{13}
- KamLand input does not improve the results on δ_{CP}
- at least $N_{\mu} m_{kt} = 2 \cdot 10^{21}$ required