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As a rapporteur from the CERN working group on Super Beams:

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# MW-linac: SPL (Superconducting Proton Linac)



### **Target and focusing system**



### Liquid target:

- Mercury
- $L = 30 \ cm$
- $R = 0.75 \ cm$
- Speed= 20 m/s

#### Magnetic Horn:

- I = 300 kA
- Toroidal magnetic field
- Charge selective

### Future:

- Optimization for the SuperBeam
- Reflector
- Mechanical/Radiation R&D

## A full beam simulation

Decay tunnel: length=20 m, radius=1 m



## SPL SuperBeam ( $\pi^+$ focused)



Flux int	ensities a	at 50	km	from	the	target
Flavour	Absolute Flux		Rel. F	lux	$\langle E_{\nu} \rangle$	
	$( u/10^{23} { m pot}/10^2 { m m}^2)$				(GeV)	
$ u_{\mu}$	1.7	$\cdot 10^{14}$		1		0.26
$\overline{ u}_{\mu}$	$4.1\cdot 10^{12}$		2.49	%	0.24	
$ u_e$	$6.1\cdot 10^{11}$		0.36	%	0.24	
$\overline{ u}_e$	$1.0\cdot 10^{10}$			0.006	s %	0.29

## SPL SuperBeam ( $\pi^-$ focused)



Flux intensities at 50 km from the target						
Flavour	Absolute Flux	Rel. Flux	$\langle E_{\nu} \rangle$			
	( $ u/10^{23} { m pot}/10^2 { m m}^2$ )		(GeV)			
$\overline{ u}_{\mu}$	$1.1\cdot 10^{14}$	1	0.23			
$ u_{\mu}$	$6.3\cdot 10^{12}$	5.7%	0.25			
$\overline{ u}_e$	$4.3\cdot 10^{11}$	3.9 %	0.25			
$ u_e$	$1.6\cdot 10^{10}$	0.15%	0.29			

$$U = \begin{pmatrix} c_{13}c_{12} & c_{13}s_{12} & s_{13}e^{-i\delta} \\ -c_{23}s_{12} - s_{13}s_{23}c_{12}e^{i\delta} & c_{23}c_{12} - s_{13}s_{23}s_{12}e^{i\delta} & c_{13}s_{23} \\ s_{23}s_{12} - s_{13}c_{23}c_{12}e^{i\delta} & -s_{23}c_{12} - s_{13}c_{23}s_{12}e^{i\delta} & c_{13}c_{13} \end{pmatrix}$$

# Measuring $heta_{13}$

$$p(\nu_{\mu} \rightarrow \nu_{e}) = -4Ue(1)Ue(2)U_{\mu}(1)U_{\mu}(2)\sin^{2}(1.27 \cdot \delta m_{12}^{2} \cdot L/E) -4Ue(1)Ue(3)U_{\mu}(1)U_{\mu}(3)\sin^{2}(1.27 \cdot \delta m_{23}^{2} \cdot L/E) -4Ue(2)Ue(3)U_{\mu}(2)U_{\mu}(3)\sin^{2}(1.27 \cdot (\delta m_{23}^{2} + \delta m_{12}^{2}) \cdot L/E)$$
(1)



M. Mezzetto, "SuperBeam studies at CERN", Nufact 01, Tsukuba, May 24-30, 2001

## **Regarding the parameters' choice.**

- Amplitude is driven by  $\sin^2(2\theta_{13})$
- Wavelength is driven by  $\delta m^2_{23}$

• But also  $\delta m^2_{12}$ , its sign,  $\delta$ ,  $\sin^2 2\theta_{23}$ ,  $\sin^2 2\theta_{12}$  have sizable effects



### $\nu_e$ Appearance Experiment

Aim of the experiment: detect  $u_e$  generated by  $u_\mu o 
u_e$  transitions with  $p(
u_\mu o 
u_e) \le 5\cdot 10^{-3}$ 

- It's a counting experiment:  $\nu$  energy is poorly reconstructed mostly because of the Fermi motion and does not allow to subdivide energy distribution in bins.
- Detector Backgrounds
  - $\mu$  mis-identified
  - $\pi^0$  production in Neutral Current
- Irreducible  $\nu_e$  beam background: slightly below  $4 \cdot 10^{-3}$ .
- Atmospheric neutrino background: negligible, thanks to the accumulator.
- **Close Detector**: A 0.5 kton detector at 1 km from the target will normalize the backgrounds with a 2% statistical accuracy.

Two reference values for L:

- The optimal one (depending from oscillation parameters and  $f_B$ ):  $L\simeq 90~{\rm km}$
- The practical one: CERN-Modane:  $L\simeq 130~{\rm km}.$

### Interesting features of a low energy conventional neutrino beam.

### $\nu$ beam:

- $\langle E_{\nu_{\mu}} \rangle \simeq 0.25 \; {\rm GeV}$
- $\nu_e$  production by kaons largely suppressed by threshold effects.

 $u_e$  in the beam come only from  $\mu$  decays.



they can be predicted from the measured  $\nu_{\mu}$  CC spec- $\Rightarrow$  trum both at the close and at the far detector with a small systematic error of  $\sim 2\%$ .

### **Detector Backgrounds**

- Good e/ $\pi^0$  separation following the large  $\pi^0 \to \gamma \gamma$  opening angle
- Good  $e/\mu$  separation in a Čerenkov detector because  $\mu$  are produced below or just above the Čerenkov threshold.
- Charm and  $\tau$  production below threshold.

# Less exiting aspects of a low energy neutrino beam

- Cross sections are small ⇒
   large detectors are necessary in spite of the very intense neutrino beam.
- $\overline{\nu}_{\mu}$  production is disfavoured for two reasons:
  - Smaller  $\pi^-$  multiplicity at the target.
  - $\overline{\nu}_{\mu} / \nu_{\mu}$  cross section ratio is at a minimum (1/5).
- Visible energy is smeared out by Fermi motion

### **Three detectors scenario**

### Water Čerenkov detector á la SuperK (40% PMT coverage) with 40 kton fiducial mass

- $\epsilon_s \sim 70\%$  (from a full simulation + analysis)
- $f_B(\pi^0 \ /e) \sim 0.002$  (full simulation + analysis using energy flow fitter to identify  $\pi^0$  )
- $f_B(\mu/e) \sim 0.001$  (full simulation)
- $f_B \,({
  m Beam}) \sim 0.003$  (full simulation of beam)

 $f_B$  are normalized to the  $u_\mu$  CC rate

# Liquid scintillator á la MiniBoone (10% PMT coverage) with 40 kton fiducial mass

- $\epsilon_s\sim 50\%$  (using MiniBoone numbers)
- $f_B(\pi^0 \ / e) \sim 0.001$  (using MiniBoone numbers)
- $f_B(\mu/e)\sim 0.001$  (using MiniBoone numbers)
- $f_B (\text{Beam}) \sim 0.003$  (full simulation of beam)

### Water Čerenkov detector á la SuperK with 400 kton fiducial mass

• Extrapolated from the 40 kton detector

## Reducing atmospheric neutrino background with the accumulator



## Water Čerenkov: Particle Identification Cut

Use Čerenkov light pattern (including opening angle, if possible) as primary  $\mu$  rejection

Tighten cut to reduce miss-ID further

 $\nu_e$  CC efficiency: 94%

 $u_{\mu}$  CC efficiency:  $\sim 1\%$ 

Muon decay identification using delayed coincidence

Only  $\sim 22\%$  of  $\mu^-$  absorbed before decay.

Visible energy cut:  ${\rm E_{vis}}>100~MeV\Rightarrow \nu_{\mu}~{\rm CC~efficiency:}\sim 0.1\%$ 



## $\pi^{\circ}$ /e background rejection based on sophisticated energy flow fitter

When the standard SuperK algorithm finds a single electron-like ring, a special algorithm is started that always identifies second candidate. Electron is identified if  $m_{\gamma\gamma} < 45 \ MeV$ .



## **Final Background Rejection**



## Final event rates for a 200 kton-year exposure

 $\nu_\mu$  CC are computed NEGLECTING  $p(\nu_\mu \to \nu_\mu$  ).  $\nu_\mu$  disappearance reduces  $\nu_\mu$  background by an additional factor 5.

Water Čerenkov, $\pi^+$ focused beam							
	Initial	Visible	Single-ring	Tight	No	$m_{\gamma\gamma} < 45$	
Channel	sample	events	100-450MeV	PID	$\mu  ightarrow$ e	$(MeV/c^2)$	
$ u_{\mu}$ CC	3250	887	578	5.5	2.5	1.5	
$ u_e{ t CC}$	18	12	8.2	8.0	8.0	7.8	
NC	2887	37	8.7	7.7	7.7	7.5	
$ u_{\mu} \rightarrow \nu_{e} $		82.4%	77.2%	76.5%	70.7%	70.5%	
Water Čerenkov, $\pi^-$ focused beam							
$ u_{\mu}$ CC	539	186	123	2.3	0.7	0.7	
$ u_e{ t CC}$	4	3.3	3.	2.7	2.7	2.7	
NC	687	11.7	3.3	3.	3.	0.3	
$\overline{ u}_{\mu} \rightarrow \overline{ u}_{e}$		79.3%	74.1%	74.0%	67.1%	67.1%	

#### **Liquid Scintillator**

	$\pi^+$ focused bear	m	$\pi^-$ focused beam			
Channel	Initial sample	Final sample	Channel	Initial sample	Final sample	
$\nu_{\mu} CC$	2538	2	$\overline{\nu}_{\mu} CC$	451	0.5	
$ u_e \ ^{CC}$	12	6	$\overline{\nu}_e \ ^{CC}$	2.3	1	
NC (visible)	48	0.5	NC	10	0.1	
$ u_{\mu} \rightarrow \nu_{e} $	100%	50%	$\overline{ u}_{\mu} \rightarrow \overline{ u}_{e}$	100%	50 %	

## **Sensitivity curves**





## $\nu_{\mu}$ disappearance experiment (liquid scintillator detector).



# Sensitivity on $\delta m^2_{23}$ and $sin^2(2\theta_{23})$

L=130



- Take three testing points at  $\sin^2 2 heta_{23} = 0.98$  ,  $\delta m_{23}^2 = 3.8, \, 3.2, \, 2.5 \, \cdot 10^{-3} \, \mathrm{eV}^2$
- 200 kton-year exposure
- 2% systematic errors
- $\sin^2 2\theta_{23}$  measured at 1%
- $\delta m^2_{23}$  measured with  $\sim 1\cdot 10^{-4}$  resolution

### A preliminary exercise on CP sensitivity

- Assume the upper value of LMA:  $\delta m^2_{12} = 10^{-4} \; eV^2$
- The CP violating observable is  $\frac{N(e^+)-N(e^-)}{N(e^+)+N(e^-)}$ , corrected for the different fluxes and cross sections. Here  $e^-(e^+)$  indicates all the e-like events selected with the  $\pi^+(\pi^-)$  focused beam.
- Run for 2 years with the  $\pi^+$  focused beam and 10 years with the  $\pi^-$  focused beam, to compensate the unfavorable ( $\overline{\nu}_e / \nu_e$ ) cross section ratio
- Fit simultaneously  $\delta$  and  $\theta_{13}$  on N(e<sup>+</sup>) and N(e<sup>-</sup>) separately.
- Take  $\theta_{13} = 5^{\circ}, 8^{\circ}, 10^{\circ}$  ( $sin^2(2\theta_{13}) = 0.03, 0.08, 0.12$ ) and a maximally violating CP phase,  $\delta = \pm 90^{\circ}$

### $\Downarrow$ (see figure)

- CP sensitivity does not worsen very much with  $\theta_{13}$ .
- In the 40 kton detector, 90% CL, a maximally violating CP phase ( $\delta = \pm 90^{\circ}$ ) would be just distinguishable from a non violating CP phase ( $\delta = 0^{\circ}$ ).
- With the 400 kton detector the prospects to observe CP violation are much improved.

## **Preliminary CP sensitivity**



## **CERN-Modane**

Two very good reasons to put the far detector deep underground:

- A detector 300 (3000) times bigger than LSND but with the same signal event rate cannot tolerate the sea-level througoing muon rate.
- It's well known that a 40 (400) kton detector deep underground would have outstanding capabilities on other themes like proton decay, atmospheric neutrinos, supernovae neutrinos, possibly solar neutrinos.

Two long highway tunnels exist at about 100 km from CERN: Monte Bianco (90 km) and Frejus (130 km), both tunnels connect Italy to France. In the Monte Bianco tunnel the Nusex experiment was performed, the underground Modane laboratory at Frejus, 5000  $m^3$ , hosts the Nemo and Edelweiss experiments.

Excaving works to dig a safety tunnel will probably start at Frejus.

The biggest cavern than can be excavated at Frejus is  $10^6 m^3$ :  $60 \times 30 \times 600 m^3$  at 2500 m.w.e. or  $50 \times 25 \times 800 m^3$  at 4400 m.w.e.

A strong interest for a SuperBeam-proton decay massive deep underground detector is fast rising in Europe, following the UNO and HyperK initiatives.

An international workshop will be organized at CERN by the end of the year: "Workshop on large detector for  $\nu$ SuperBeam, p decay and atmospheric  $\nu$ 's", jointly with the UNO collaboration and hopefully with HyperK.

## **Conclusions**

- A full simulation of the beam line and neutrino events together with real analysis of the events have been performed.
- Low energy is the best option for SuperBeams as far as concern backgrounds.
- Very MASSIVE detectors are needed
- Čerenkov detectors guarantee the necessary performances ⇒ SuperBeam + proton decay+ atmospheric and supernovae neutrinos at the same time.

A 40 kton (fiducial) detector can:

- Detect  $\sin^2(2\theta_{13})$  if  $\sin^2(2\theta_{13}) \ge 0.007 (90\% CL)$
- Measure  $\sin^2(2 heta_{23})$  with 1-2% resolution
- Measure  $\delta m^2_{23}$  with  $10^{-4} \, eV^2$  resolution.
- Detect CP violating effects at 90%CL in a small region of parameters space.

A 400 kton (fiducial) detector can:

- Detect  $\sin^2(2\theta_{13})$  if  $\sin^2(2\theta_{13}) \ge 10^{-3} (90\% CL)$
- Detect CP violating effects at more than  $3\sigma$  in a wider range of parameters space.