MACRO results on v oscillations

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#### for the MACRO Collaboration





Les Houches, 18-22 June 2001 - Neutrino Masses and Mixings

## Outline

- Neutrino detection in MACRO @ Gran Sasso updated to acq end (19 Dec 2000)
- Results on through-going upward muons: matter effects  $\nu\mu \rightarrow \nu\tau$  or  $\nu\mu \rightarrow \nu\tau$
- Low energy topologies
- Multiple scattering

Different technique than SK clear event topologies, different energies, fully automatic analysis, no rejection of events at trigger level (efficiency, acceptance, backgrounds can be studied using atm. muons)

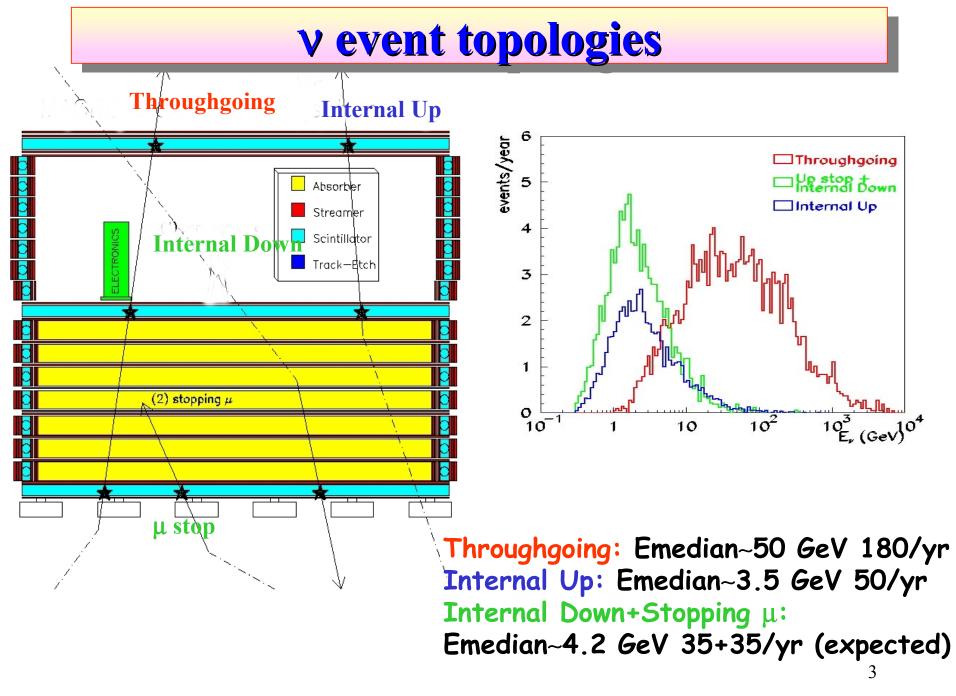
- Min rock coverage 3150 hg/cm<sup>2</sup>
- 600 ton liquid scintillator (600 ps), 20000 m<sup>2</sup> streamer tubes (< 1°)</li>

76.6×12×12 m<sup>3</sup>

E<sub>th</sub>~1 GeV

*(a)*vertical

NIM A324(1993)



### **Upward throughgoing µ analysis**

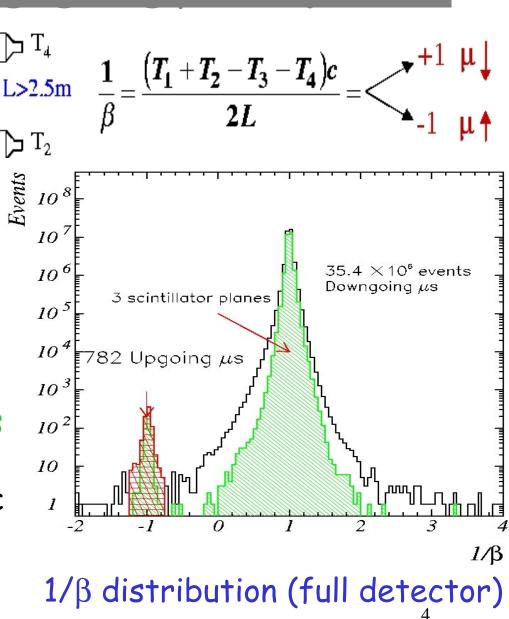
Main cut: position along  $T_1 \leftarrow T_1 \leftarrow T_2 \leftarrow T_1 \leftarrow T_2 \leftarrow T_$ 

T.o.F. technique:  $T_3 \subset$ 

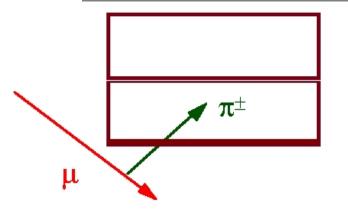
~200 gr/cm<sup>2</sup> in rock Absorber to reduce at 1% Background from upgoing  $\pi s$ 

No scanning, fully automatic

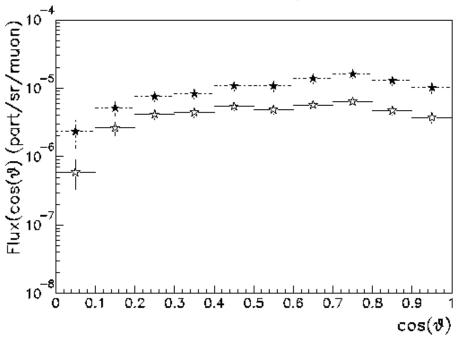
-1.25 < 1/β < -0.75

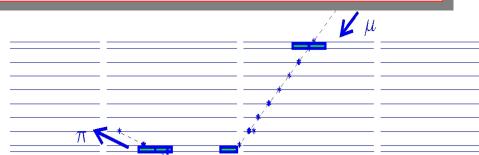


## The backgrounds



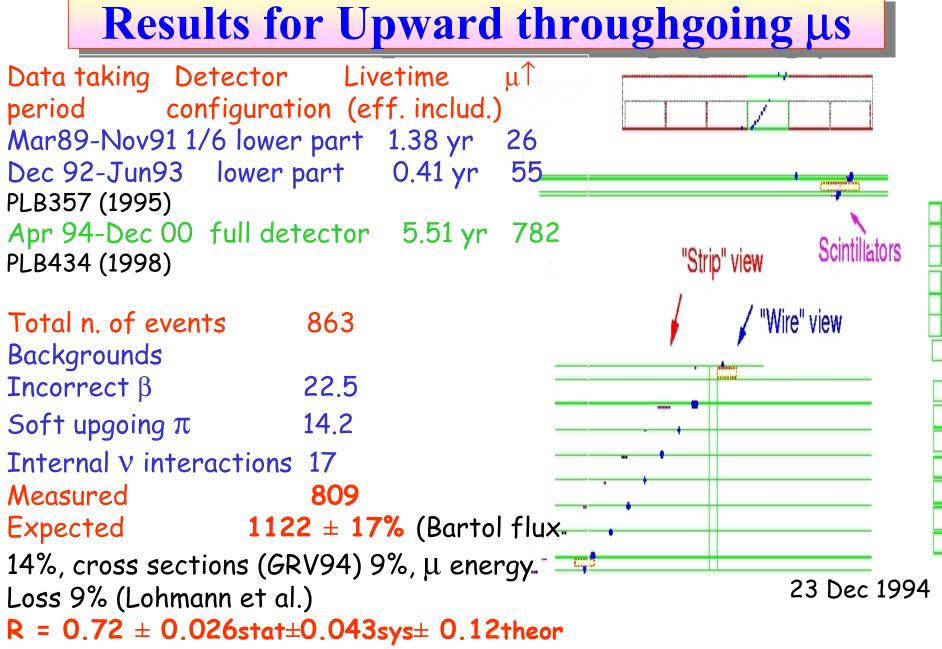
MACRO Coll., Astr. Phys 9 (1998)



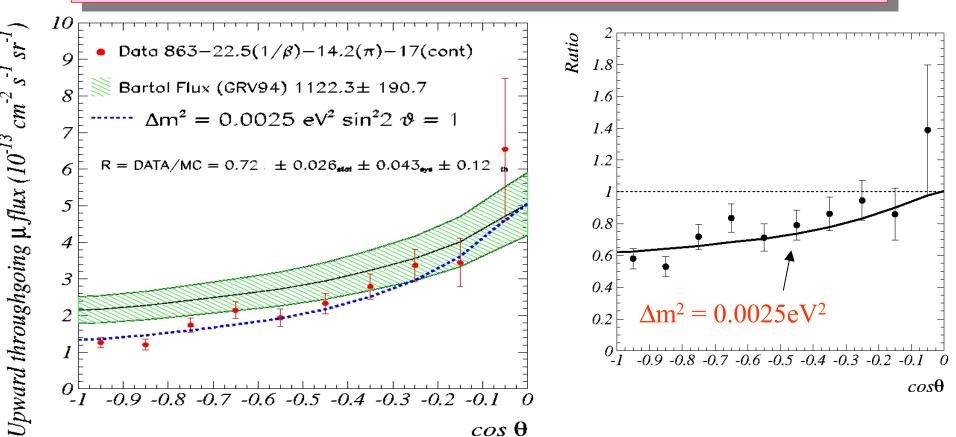


RUN = 9967 EVENT= 3941 11-MAY-95 07:36:38

Physical background from atm.  $\mu \downarrow$  photonuclear interactions producing upgoing soft particles Important for shallow detectors (Baksan, IMB while SK and Soudan2 have vetos) 243 upgoing particles between 12.2  $\cdot 10^6 \ \mu s \downarrow \Rightarrow \sim 10^{-4} \ \pi/\mu \downarrow$ ~1% in throughgoing  $\mu$ s<sup>↑</sup> ~5% in stopping  $\mu$ s <sup>5</sup>

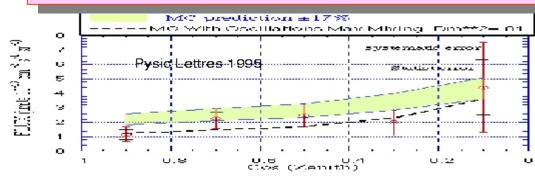


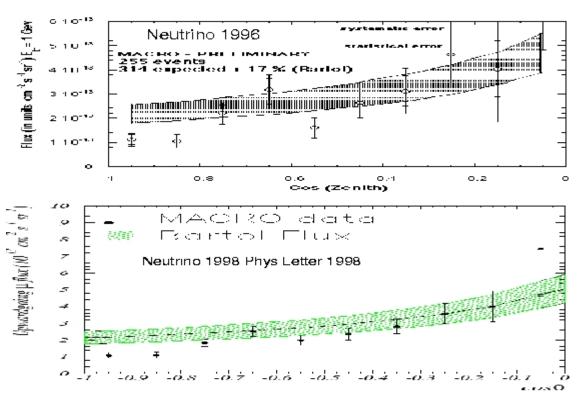
## $\mu\uparrow$ flux angular distribution



 $\chi^2$  test on the angular distribution (10 bins) with prediction normalized to data:  $\chi^2/dof=25.9/9$  for no-oscillations  $\Rightarrow P = 0.2\%$  $\chi^2/dof=9.6/9$  for  $\nu\mu \rightarrow \nu\tau$  ( $\Delta m^2 = 0.0025 \text{ eV}^2 \sin^2 2\theta = 1$ )  $\Rightarrow P = 37\%$ 

## $\mu^{\uparrow}$ flux angular distribution during time

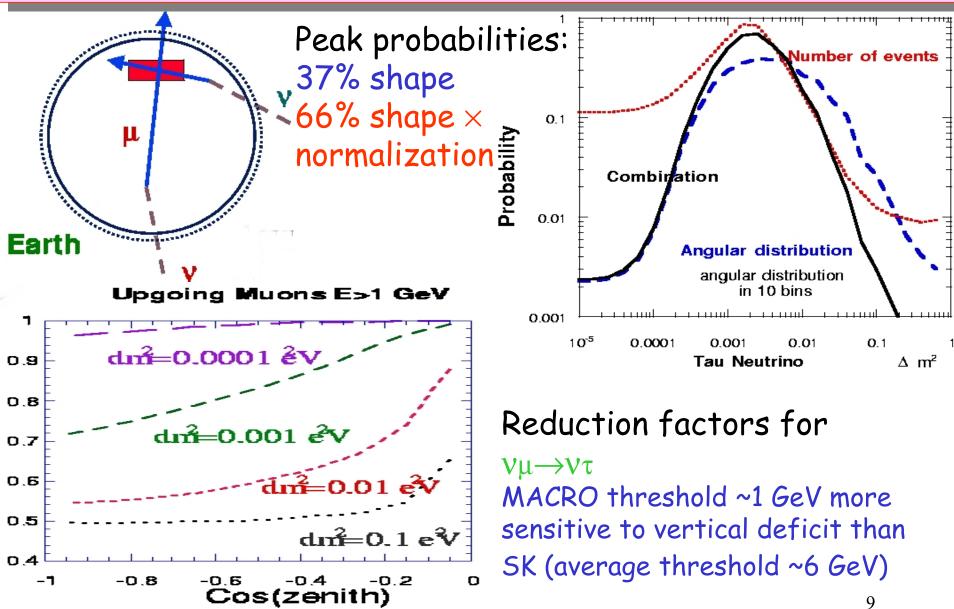




Since 1994 to now the angular distribution has become more regular

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#### $\nu\mu \rightarrow \nu\tau$ oscillations in upward throughgoing $\mu$



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### **Angular distribution for selected sample**

 $\cos \theta$ 

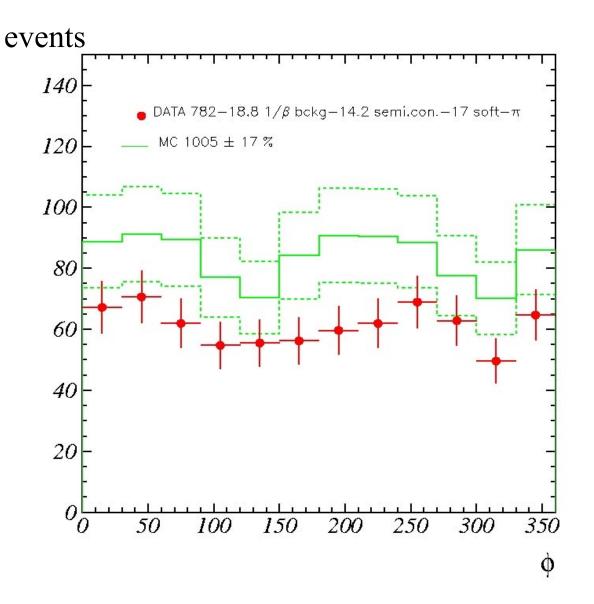
10Data  $430 - 1.25(1/\beta) - 9.6(cont)$ 💓 Bartol Flux (GRV94) 630.4± 107.2  $cm^{-2}$ •••••  $\Delta m^2 = 0.0025 \text{ eV}^2 \sin^2 2 \vartheta = 1$ Jpward throughgoing  $\mu$  flux (10<sup>-13</sup>  $R = DATA/MC = 0.665 \pm 0.032_{stat} \pm 0.040_{svs} \pm 0.113_{th}$ 6 4 3 2 -0.9 -0.8 -0.7 -0.6 -0.5 -0.4 -0.3 -0.2 -0.1 0

3 box events (redundant time measurement) Same shape of full sample

 $\chi^{2}$ /dof=9.4/7 for no-oscillations ⇒ P = 22.8%

 $\chi^{2}$ /dof=3.7/7 for  $\nu\mu \rightarrow \nu\tau$  $(\Delta m^2 = 0.0025 \text{ eV}^2)$ sin<sup>2</sup>2θ=1) ⇒ P = 81.5%10

### Angular distribution for selected sample

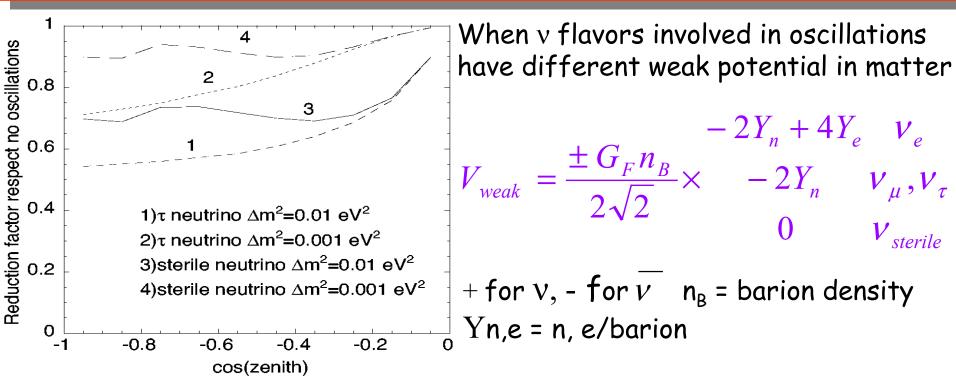


Oscillations do not affect azimuth angular distribution

 $\chi^2/dof=2.8/11$ normalizing prediction to data  $\Rightarrow$ P = 0.99

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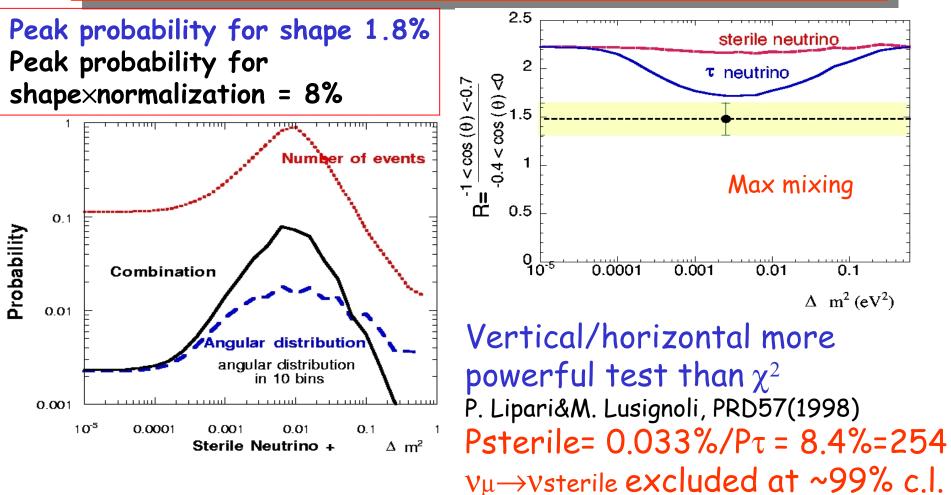
## **Matter effects in upward throughgoing µs**



Matter effects can be important for  $\nu\mu \rightarrow \nu e$ ,  $\nu sterile$  not for  $\nu\mu \rightarrow \nu \tau$  $E_{v/\Delta m^{2}} \ge 10^{3} GeV/eV^{2} \Rightarrow$  for HE up-throughgoing  $\mu s$ 

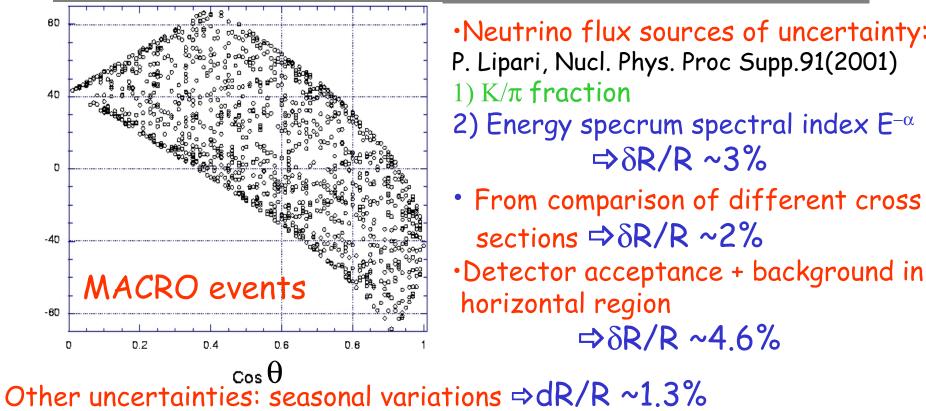
For maximal mixing matter effect produces a reduction of oscillation effect ⇒ closer to predicted with no oscillations For mixing <1 enhancement for some values of parameters (MSW)





Ratio: sensitive to sign of deviation, gain in statistical significance (2bins) but some feature of angular shape could be lost MACRO bin choice through simulation (MACRO Coll., hep-ex/0106049)

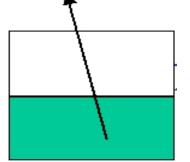
## Systematic errors in the ratio



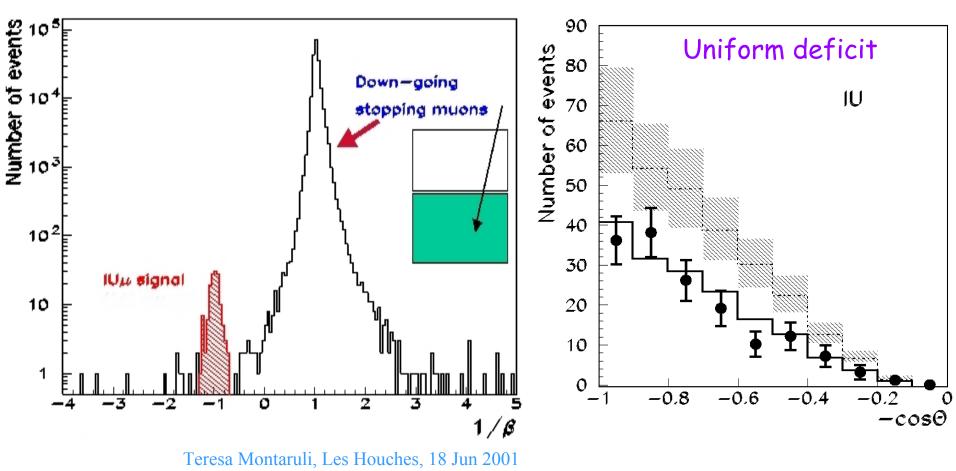
.atitudine

Other uncertainties: seasonal variations ⇒dR/R ~1.3% + US atmosphere profile ⇒dR/R ≤1% MACRO upgoing µs: variation in vertical/horizontal ratio R(Nov-Apr) - R(May-Oct)=≈0.19±0.17stat Seasonal effect difficult to calculate because neutrinos come from meson and muon decays and from all over the Earth

## **Internal Up events**



ToF between central and upper scintillator layers+ containment cuts to reject up-throughgoing  $\mu$ s (1% backg) DATA 161 with -1.3< 1/ $\beta$  <-0.7 (eff. livetime 0.58yr) Backgrounds (wrong  $\beta$ , secondary hits) = 7  $\Rightarrow$  154

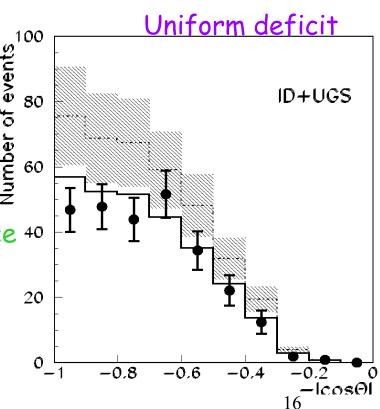


# **Internal Down + Upward Stopping**

Selected using topological cuts Tracks inside fiducial volume crossing bottom scintillator layer ending in lower part, visual scanning (real+simulated events),  $\geq 100 \text{ gr/cm}^2$ : background due to soft  $\pi$ s (~5%) DATA: 272 events (5.6 yr), background 10 events  $\Rightarrow 262$ 

Predictions:  $\Phi = \Phi_V \otimes \sigma_V \otimes \epsilon(E\mu,\theta)$   $\Phi_V = Bartol v flux (~20\%)$   $\sigma_V = Qel+1\pi+DIS(GRV94-LO) (~15\%)$ Lipari et al., PRL74(1995) 25% error in MC normalization  $\epsilon(E\mu,\theta) = detector responce and acceptance 40$ 

(~10%)
6% difference respect to MINOS Neugen
IU and ID+UGS mainly due to CCVμ
NC and ve (~13% IU and 10% UGS+ID)



#### **Double Ratio**

#### IU

#### DATA

154 ± 12stat

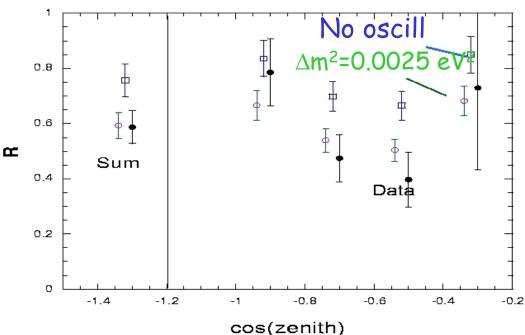
MC (no osc) $285 \pm 28sys \pm 71th$ R=DATA/MC(no osc) $0.54 \pm 0.13tot$ MC ( $\Delta m^2$ =0.0025 eV²)168

ID+UGS 262 ±16stat

375 ±37*s*ys ±94th 0.70 ± 0.19tot 284

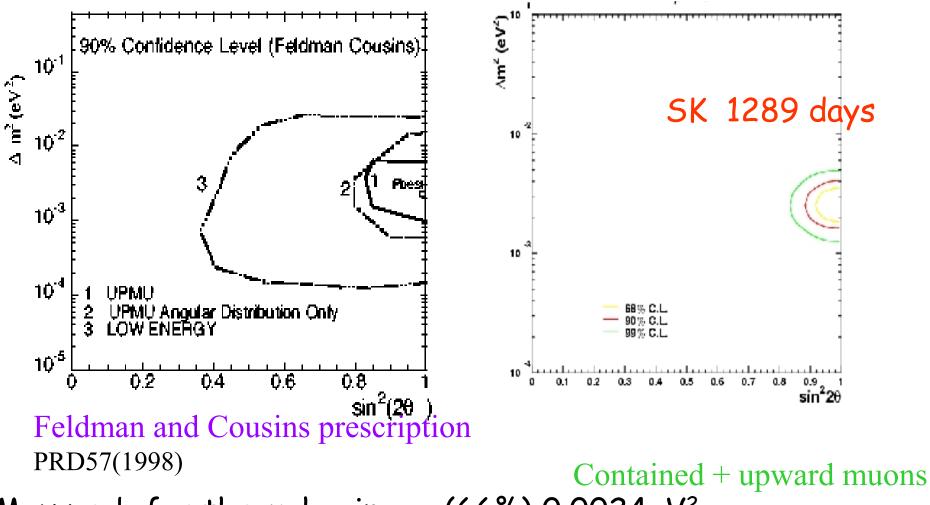
 $\begin{array}{l} R = IU/(ID + UGS)_{DATA} = 0.59 \pm \ 0.060 \text{stat} \ R = IU/(ID + UGS)_{MC} = 0.76 \pm \ 0.059 \text{sys} \\ R = IU/(ID + UGS)_{MC} (\Delta m^2 = 0.0025 \ \text{eV}^2) = 0.59 \pm \ 0.046 \text{sys} \end{array}$ 

R<sub>IU</sub> and R<sub>ID+UGS</sub> not same reduction ⇒deficit not due to theor. overestimate



of v flux/cross sections (both topologies <Ev>~4 GeV) Expected reductions for  $\Delta m^2 \sim 1-10 \times 10^{-3} eV^2 \sin^2 2\theta = 1$ 1/2 for IU 1/4 for ID+UGS Probability to obtain double ratio so far from expected is 2.2% (including non-gaussian shape of the uncertainty of the ratio)





Max prob for throughgoing  $\mu s$  (66%) 0.0024eV^2 for low energy (87%)

## **Multiple Scattering**

- Upgoing muon energy estimate through multiple scattering in 25 radiation lengths of MACRO using ST system ⇒ L/E 2 analysis:
- 1) Angular and space difference between lower track (1-8 planes) upper track (6:14)
- MACRO ST 3 cm cells (digital information)⇒
- $\sigma \sim 3 \text{cm} \sqrt{12} \sim 1 \text{cm} \Rightarrow \text{residual E} \mu 10 \text{GeV} / \sqrt{\cos \vartheta}$

2) ST Space resolution improved by a factor 3.5 to 3mm using TDC analogue information (150 ns)  $\Rightarrow$  muon residual energy up to 40 GeV (Ev~200 GeV) Given preferred oscillation parameter  $\Delta m^2 ~O(10^{-3} \text{ eV}^2)$  oscillation effects should be stronger at Eµ≤10 GeV and disappear at E> 100 GeV Results (positive) will be presented at ICRC

## Conclusions

1) High energy events: angular distribution more regular than in the past χ²/dof=9.6/9 for νµ→ντ (Δm² = 0.0025 eV² sin²2θ=1) ⇒P = 37% (only from shape) 66% (shape + normalization) Two flavor sterile v oscillations disfavored respect to νµ→ντ @ 99% c.l. for max mixing

2) Low energy events: flat angular distribution deficit, up/down asymmetry P~2% (no oscillation) independent on theoretical errors Confirm HE μs preferred parameters

3) Multiple scattering analysis provides residual  $\mu$  energy information and L/E indication