

Leptogenesis in left-right symmetric theories

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↳ masses and mixing, Les Houches
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- Motivation
- Model + Strategy
- Results on χ_B
- ~~CP~~ phases and $O_\nu \beta \beta$
- Conclusion + Outlook

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hep-ph/0704228

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MOTIVATION

- Baryogenesis in SM does not work
- Neutrino oscillation needs non SM physics



Model predicts heavy neutrinos

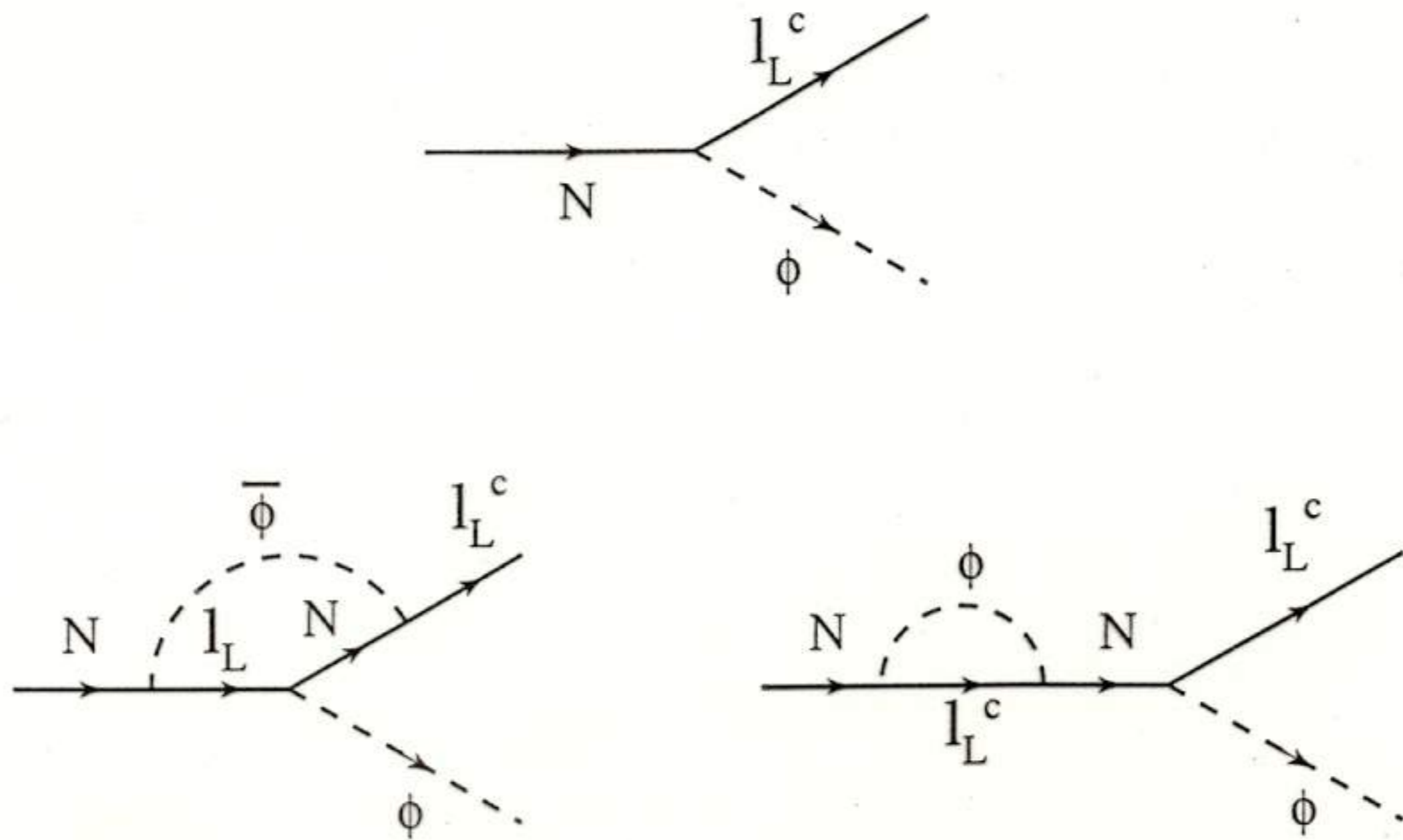
- Leptogenesis explains $\eta_B \sim 10^{-10}$ through decay of heavy neutrinos

- ? Does your model for neutrino mass ?
- also produce correct baryon asymmetry •

Ellis, Lola, Nanopoulos 99 / Berger + Brahmachari 99 /
Lazarides, Vladimos 99 / Barbieri et al 99 / Berger 00 /
Buchmüller, Plumacher 99 / Jeanrenaud, Kraljic, Lazarides 00 /
Kang, Kang, Sarkar 00 / Goldberg 00 / Nezi + Orloff 00
Falcone, Tramontano 01 / Nilsson, Takahashi 01 /
Buchmüller + Yanagida 98 / Hirsch + King 01

Leptogenesis à la Fukugita + Yanagida:

Radiative corrections create lepton asymmetry



$$\epsilon_1 = \frac{\Gamma(N_1 \rightarrow l^c \phi) - \Gamma(N_1 \rightarrow l \phi^\dagger)}{+}$$

$$= \frac{1}{8\pi v^2} \frac{1}{(m_D^\dagger m_D)_{11}} \sum_{j=2,3} \text{Im} \left\{ (m_D^\dagger m_D)_{1j}^2 \right\} f \left(M_j^2 / M_1^2 \right)$$

where $f \left(M_j^2 / M_1^2 \right) \approx \frac{3}{2} \frac{M_1}{M_j}$ for $M_j^2 \gg M_1^2$

from coupling:

$$-\mathcal{L} = \overline{l_{\alpha L}} \frac{\phi}{\langle \phi \rangle} m_{D\alpha\beta} N_{\beta R}$$

$$Y_B = a Y_L = a \frac{1}{g^*} \kappa \epsilon_1 \quad \text{where}$$

$$a = \frac{c}{c-1} \approx -0.55$$

$$c = \frac{8N_f + 4N_H}{22N_f + 13N_H} \approx 0.35 \quad \text{for } \begin{matrix} N_f = 3 \\ N_H = 1, 2 \end{matrix}$$

Fraction of lepton asymmetry Y_L converted into baryon asymmetry Y_B (sphaleron)

$$g^* \approx 110$$

massless degrees of freedom @ $T = M_1$

κ : Dilution factor

$$l^c \phi \rightarrow l \phi^\dagger \longleftrightarrow N_1 \rightarrow l^c \phi$$

$$\kappa \approx \begin{cases} \sqrt{0.1k} \exp\left\{-\frac{4}{3} \sqrt{0.1k}\right\} & k \geq 10^6 \\ \frac{0.3}{k(\ln k)^{0.5}} & 10 \leq k \leq 10^6 \\ \frac{1}{2\sqrt{k^2+9}} & 0 \leq k \leq 10 \end{cases}$$

$$k = \frac{\Gamma_1}{H(T=M_1)} = \frac{(m_0^+ m_0)_1 M_1}{8\pi v^2} \frac{M_{Pl}}{1.66 \sqrt{g^*} M_1^2} \leq 10 \quad (3)$$

Leptogenesis and Neutrino mixing

$$1) m_D = m_L - \tilde{m}_D M_R^{-1} \tilde{m}_D^T \\ \simeq -\tilde{m}_D^T M_R^{-1} \tilde{m}_D \quad (\text{in most works})$$

left-handed triplet

neglected ! ?

2) m_D light left-handed \leftrightarrow Oscillation

M_R heavy right-handed \leftrightarrow Leptogenesis

Is there a connection?

→ Left-Right Symmetry

$$SU(2)_L \times SU(2)_R \times U(1)_{B-L}$$

$$\Psi_L : (2, 1, -1)$$

lepton doublet

$$\Psi_R : (1, 2, -1)$$

lepton doublet

$$\Phi : (2, 2, 0)$$

Higgs bi-doublet \tilde{m}_0

$$\Delta_R : (1, 3, 2)$$

Higgs triplet M_R

$$\Delta_L : (3, 1, 2)$$

Higgs triplet m_L

$$- \mathcal{L} = \overline{\Psi_{\alpha L}} h_{\alpha\beta} \Phi \Psi_{\beta R} + f_{\alpha\beta} \left[\overline{\Psi_{\alpha L}} \varepsilon \Delta_L \vec{\tau} \Psi_{\beta L} + (L \leftrightarrow R) \right]$$

SSB: $\langle \Phi \rangle = \begin{pmatrix} x & 0 \\ 0 & x' \end{pmatrix}$ $\langle \Delta_{L,R} \rangle = \begin{pmatrix} 0 & 0 \\ v_{L,R} & 0 \end{pmatrix}$

$$x^2 + x'^2 = v^2 \quad \text{and} \quad \frac{x}{x'} \approx \frac{m_t}{m_b} \rightarrow x \approx v$$

Mass matrix $\begin{pmatrix} v_L \not{t} & v \not{h} \\ v \not{h}^T & v_R \not{t} \end{pmatrix}$ See-saw: $v_R \gg v_L, v$

In general: $v_L v_R = \delta M_W^2$ where $\delta = \mathcal{O}(1)$

Strategy:

$$m_D = m_L - \tilde{m}_0 M_R^{-1} \tilde{m}_0^T$$



$$V_L f = m_L \quad V_R f = M_R$$

$$V_L V_R = \delta M_W^2$$

$$m_D = f \left(V_L - \tilde{m}_0 \frac{f^{-1}}{\delta M_W^2} \tilde{m}_0^T \right)$$



• LMA or SMA or QVO

• hierarchical m_i

• $\tilde{m}_0 = m_{up}, m_{down}, M_{lepton}$

$$V_R f = M_R$$



Diagonalize

Y_B

Free parameters: $\alpha, \beta \longrightarrow (m_{\nu})_{ee} : \mathcal{O}(\beta\beta)$
 $\delta \longrightarrow \text{CP} : \text{Oscillation}$
 (m_1)

"type II see-saw":

$$m_\nu = m_L - \tilde{m}_0 M_R^{-1} \tilde{m}_0^T$$

$$R_{\max} = \frac{|\tilde{m}_0 M_R^{-1} \tilde{m}_0^T|}{|m_L|} = \frac{m^2/v_R}{v_L} = \frac{m^2}{8M_W^2}$$

$$\approx \begin{cases} \mathcal{O}(1) \\ 0 \end{cases}$$

for $\tilde{m}_0 = m_{\text{up}}$

for $\tilde{m}_0 = m_{\text{down}}, m_{\text{lepton}}$

① $\tilde{m}_0 = m_{\text{up}}$ lep+ph/0704228
SO(10)

$$M_R = \frac{v_R}{v_L} \overline{m}_\nu$$

$$\text{where } (\overline{m}_\nu)_{ij} = \begin{cases} (m_\nu)_{ij} & (i,j) \neq (3,3) \\ (m_\nu)_{33} + S & i=j=3 \end{cases}$$

$$S \sim \frac{M_W m_t}{v_R} \sim 10^{-2} \text{ eV}$$

② $\tilde{m}_0 = m_{\text{down}}, m_{\text{lepton}}$ lep-ph/0705175
SU(5)

$$M_R = \frac{v_R}{v_L} m_\nu$$

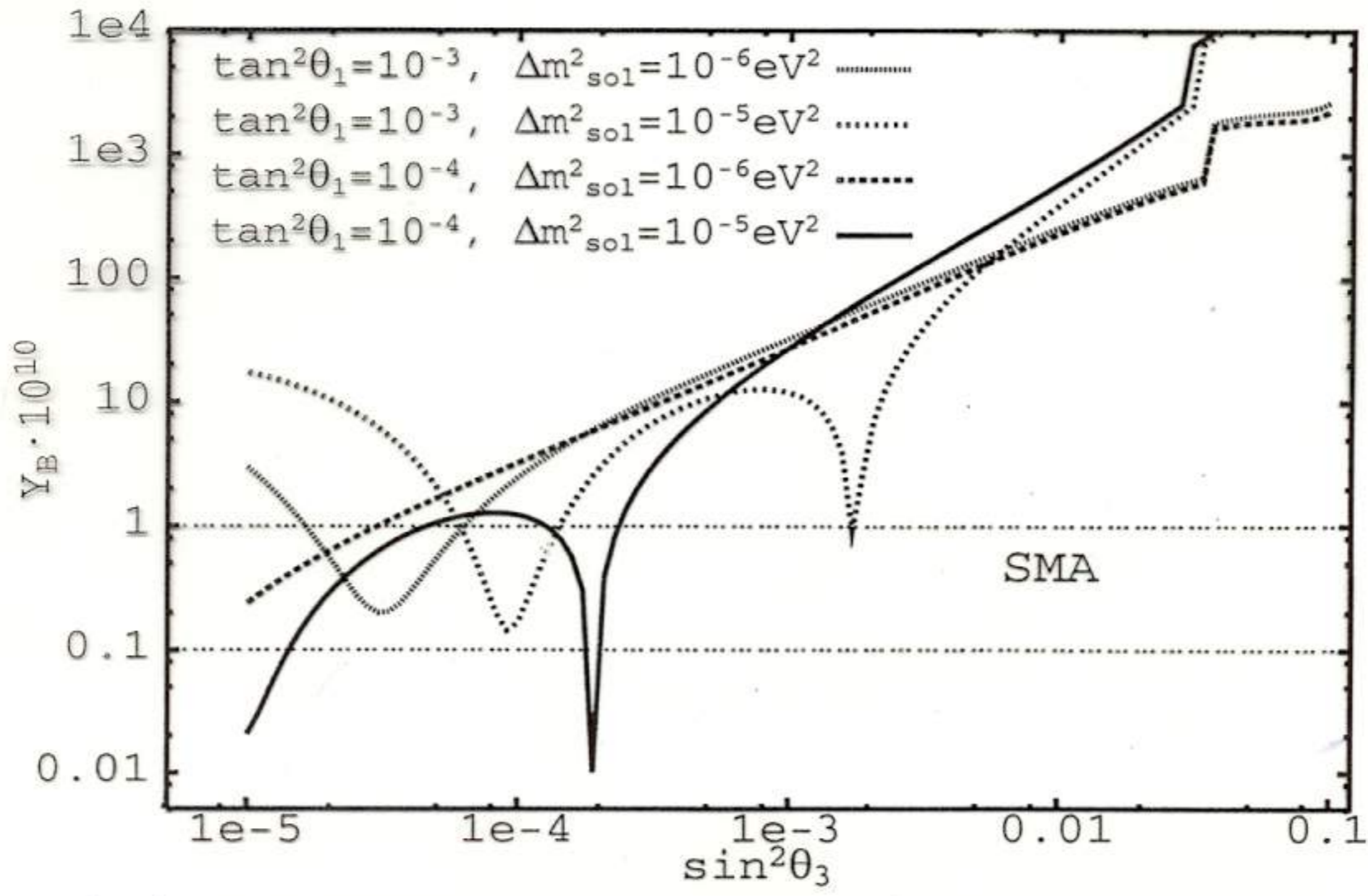
J.L. m_L dominates m_ν

①

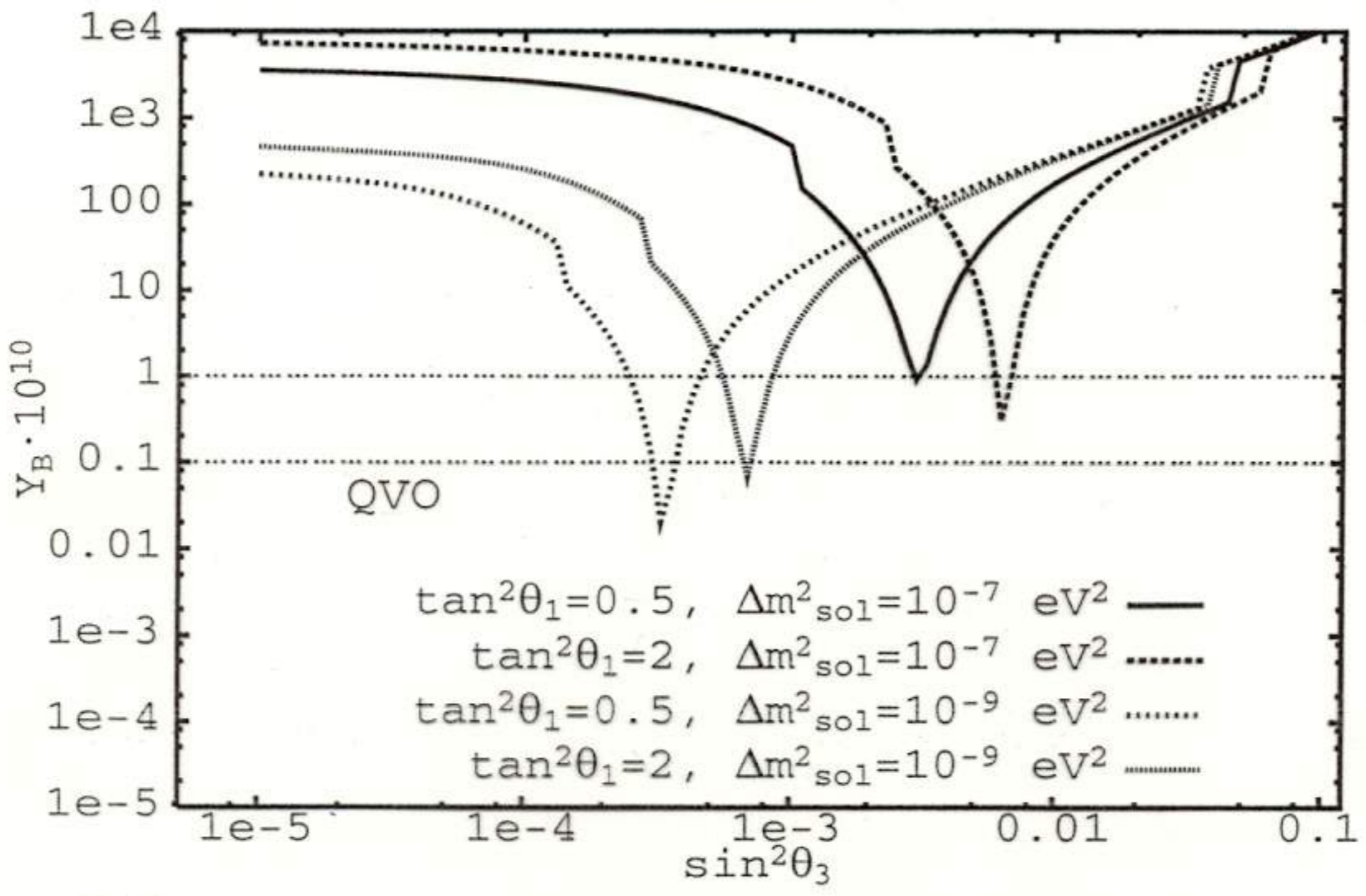
HIERARCHICAL SCHEME

DIRAC = UP quark

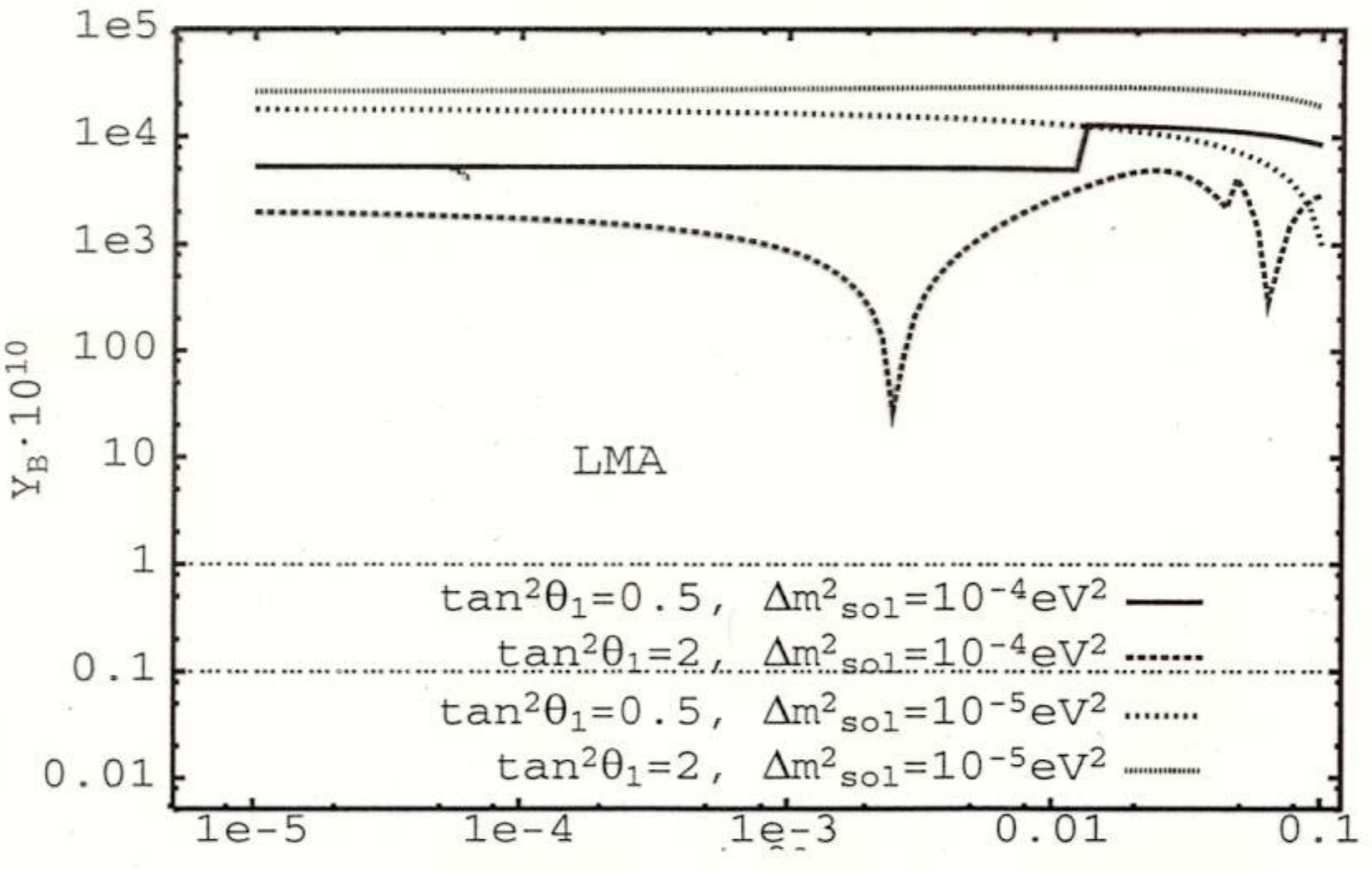
$3\alpha = 4\beta$
 $= 6\delta = \pi$



$5\alpha = 4\beta$
 $= 6\delta = \pi$



$5\alpha = 6\beta$
 $= 3\delta = \pi$



$\tilde{m}_0 = m_u \approx \text{diag}(0, 0, m_t)$

$V_R = 10^{15} \text{ GeV}$

$\delta = 1$

②

② Dirac = down quark
charged lepton

~~Majorana~~ $M_R = \frac{v_R}{v_L} m,$

$$U_L = \begin{pmatrix} c_1 c_3 & s_1 c_3 & s_3 e^{-i\delta} \\ -s_1 c_2 - c_1 s_2 s_3 e^{i\delta} & c_1 c_2 - s_1 s_2 s_3 e^{i\delta} & s_2 c_3 \\ s_1 s_2 - c_1 c_2 s_3 e^{i\delta} & -c_1 s_2 - s_1 c_2 s_3 e^{i\delta} & c_2 c_3 \end{pmatrix} P$$

$$P = \text{diag}(1, e^{i\alpha}, e^{i(\beta+\delta)})$$

→ $P(\nu_\alpha \rightarrow \nu_\beta) = P(\nu_\alpha \rightarrow \nu_\beta, \delta)$

$$\langle m \rangle = \langle m \rangle (\alpha, \beta)$$

$$U_L \approx \begin{pmatrix} 1/\sqrt{2} & 1/\sqrt{2} & s_3 e^{-i\delta} \\ -\frac{1}{2}(1+s_3 e^{i\delta}) & \frac{1}{2}(1-s_3 e^{i\delta}) & 1/\sqrt{2} \\ \frac{1}{2}(1-s_3 e^{i\delta}) & -\frac{1}{2}(1+s_3 e^{i\delta}) & 1/\sqrt{2} \end{pmatrix} P$$

for LMA, QVO

$$U_L \approx \begin{pmatrix} 1 & 0 & s_3 e^{-i\delta} \\ -\frac{1}{2}s_3 e^{i\delta} & 1/\sqrt{2} & 1/\sqrt{2} \\ -\frac{1}{2}s_3 e^{i\delta} & -1/\sqrt{2} & 1/\sqrt{2} \end{pmatrix} P$$

for SMA

With approximate forms of $U_L = U_R$:

$$K \approx 1.5 \left(\frac{10^{-6} \text{ eV}}{m_1} \right) \left(\frac{10^{15} \text{ GeV}}{V_R} \right)^2 \delta \left(\frac{m}{\text{GeV}} \right)^2 \begin{cases} 1 - 2s_3 c_\delta & \text{LMA} \\ & \text{QVO} \\ s_3^2 & \text{SMA} \end{cases}$$

Fulfilled for typical parameters; use $\delta \lesssim 1/6$:

$$Y_B \cdot 10^{10} \lesssim 4.1 \left(\frac{m}{\text{GeV}} \right)^2 \frac{1}{1 - 2s_3 c_\delta} \left\{ (s_{2\alpha} + 4s_3 s_\delta c_{2\alpha}) \frac{m_1}{\sqrt{\Delta m_{\odot}^2}} + 2(s_{2(\beta+\delta)} - 2s_3 s_{2\beta+\delta}) \frac{m_1}{\sqrt{\Delta m_A^2}} \right\}$$

LMA
QVO

$$Y_B \cdot 10^{10} \lesssim 8.2 \left(\frac{m}{\text{GeV}} \right)^2 \left\{ s_{2(\alpha-\delta)} \frac{m_1}{\sqrt{\Delta m_{\odot}^2}} + s_{2\beta} \frac{m_1}{\sqrt{\Delta m_A^2}} \right\}$$

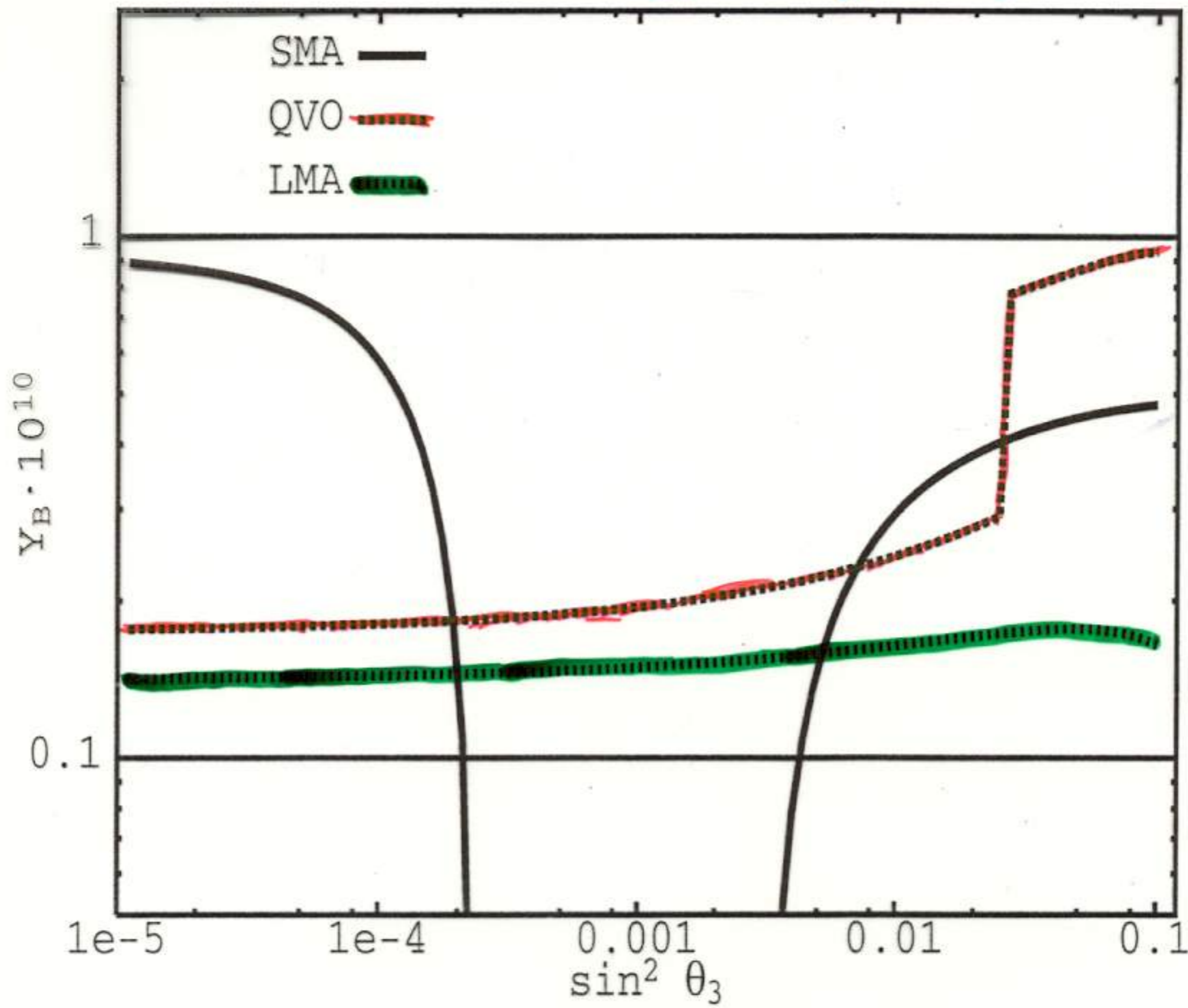
S
M
A

■ Term proportional to $\sqrt{\frac{1}{\Delta m_{\odot}^2}}$ contributes most

■ $Y_B \propto m^2 \cdot m_1$

■ If $\alpha = \beta = \delta = 0$: $Y_B = 0$

② DIRAC = down quark charged lepton



$$\Delta m_A^2 = 3 \cdot 10^{-3} \text{ eV}^2, \quad \Theta_2 = \pi/4$$

$$\Delta m_\Theta^2 = 5 \cdot 10^{-6} \text{ eV}^2, \quad \tan^2 \Theta_1 = 5 \cdot 10^{-4}, \quad m_1 = 10^{-5} \text{ eV}$$

$$\Delta m_\Theta^2 = 5 \cdot 10^{-5} \text{ eV}^2, \quad \tan^2 \Theta_1 = 1, \quad m_1 = 10^{-5} \text{ eV}$$

$$\Delta m_\Theta^2 = 10^{-8} \text{ eV}^2, \quad \tan^2 \Theta_1 = 1, \quad m_1 = 2 \cdot 10^{-6} \text{ eV}$$

$$3\alpha = 5\beta = 6\delta = \pi$$

$$\tilde{m}_D = \begin{pmatrix} 0 & \sqrt{m_d m_s} & 0 \\ \sqrt{m_d m_s} & m_s & \sqrt{m_b m_s} \\ 0 & \sqrt{m_b m_s} & m_b \end{pmatrix}$$

Maximal value for Y_B :

$$0.1 \leq Y_B \cdot 10^{10} \lesssim \begin{cases} 0.6 \left(\frac{m_1}{10^{-5} \text{ eV}} \right) & \text{SMA} \\ 1.4 \left(\frac{m_1}{10^{-6} \text{ eV}} \right) & \text{LMA} \\ 2.1 \left(\frac{m_1}{10^{-7} \text{ eV}} \right) & \text{QVO} \end{cases}$$

 m_1 not smaller than

10^{-6} eV	for	SMA
10^{-7} eV	for	LMA
10^{-8} eV	for	QVO

Lower
limit

Connection to $0\nu\beta\beta$

archical scheme has highest $\langle m \rangle$ for LMA

• High $\langle m \rangle = \sum_i U_{Lei}^2 m_i$ and correct γ_B ?

→ Random scan: $\Delta m_{21}^2 = 10^{-5} \dots 10^{-3} \text{ eV}^2$

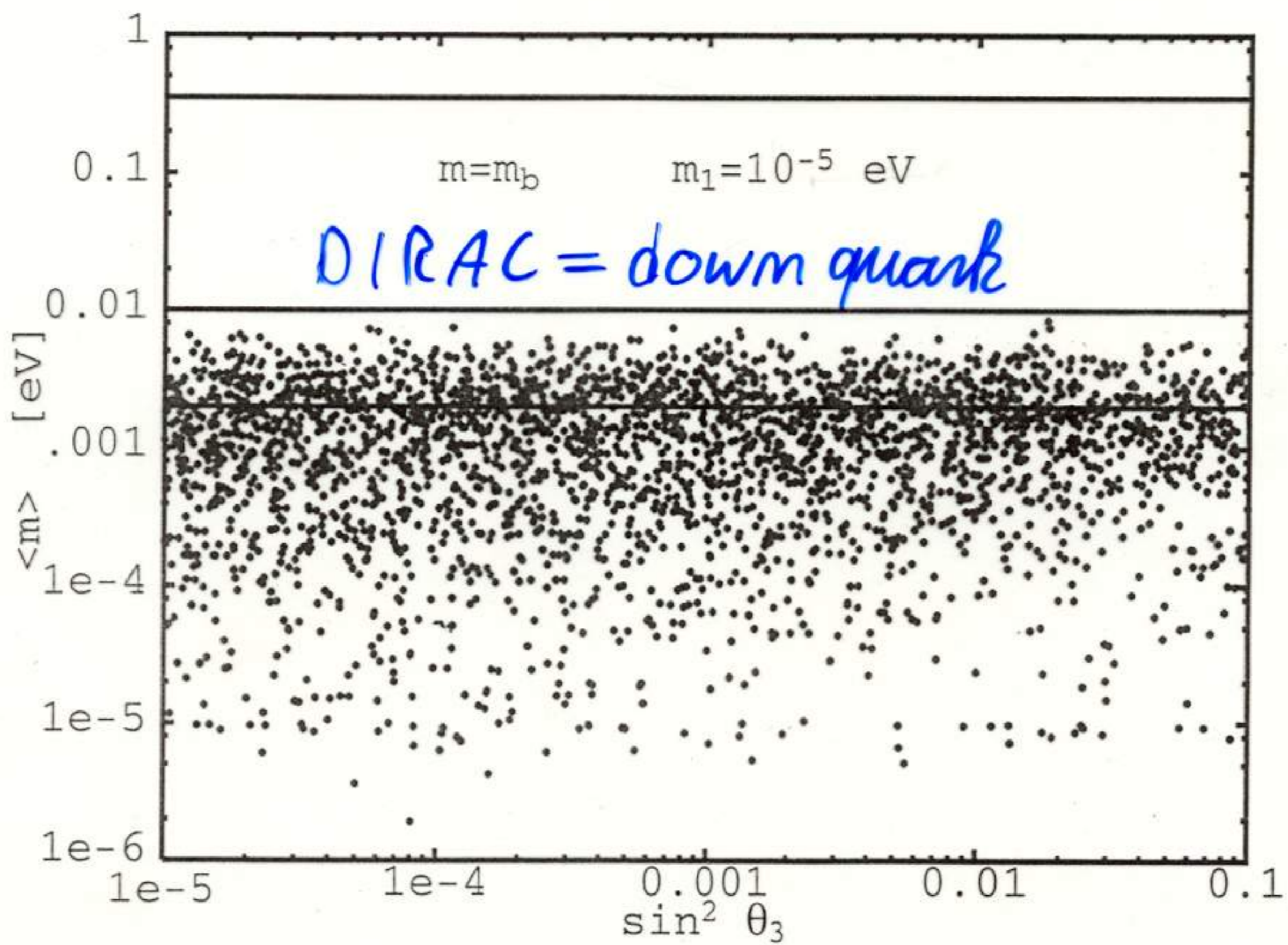
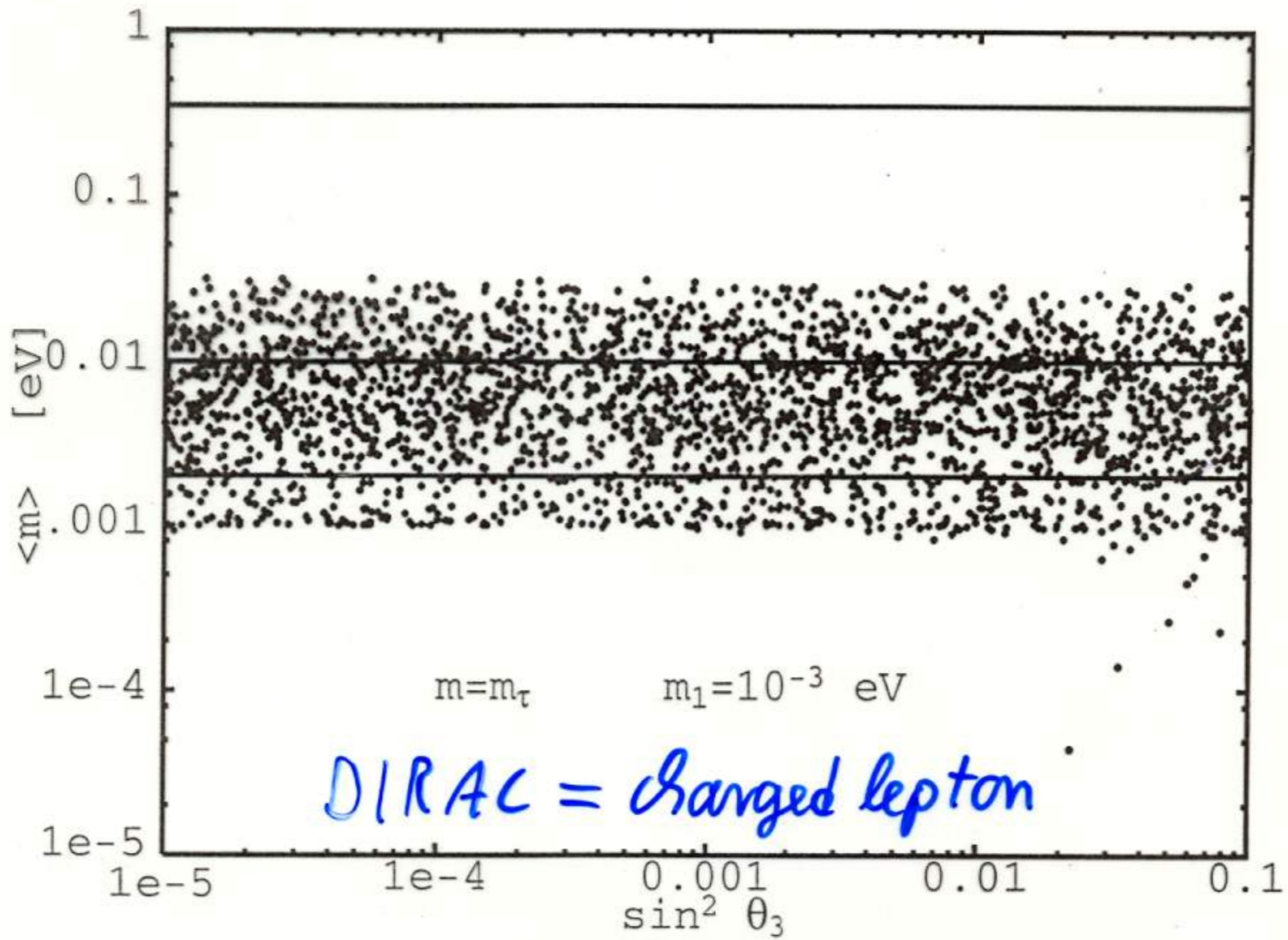
$$\tan^2 \Theta_{12} = 0.1 \dots 4$$

$$\alpha, \beta, \delta = 0, \dots, 2\pi$$

	correct γ_B	$\langle m \rangle \geq 2 \cdot 10^{-3} \text{ eV}$
$\rho_{\text{WM}} = 10^{-3} \text{ eV}$	5%	4%
$\rho_{\text{WM}} = 10^{-5} \text{ eV}$	27%	7%
$\rho_{\text{PMN}} = 10^{-3} \text{ eV}$	35%	30%
$\rho_{\text{lepton}} = 10^{-5} \text{ eV}$	1%	$\leq 1\%$

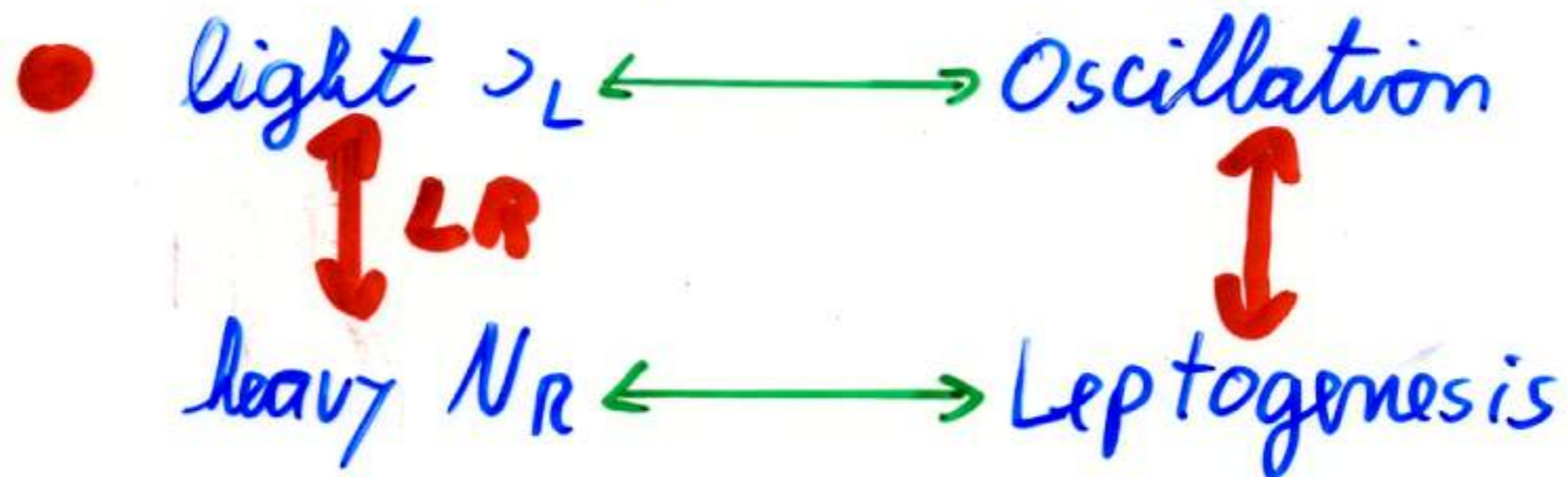
$\rho_{\text{hierarchical}}$ is equally distributed ($s_3 e^{-i\delta}$)

Parameter sets giving correct Y_B



Summary

● LR symmetry + Leptogenesis



● type II see-saw, m_L dominates

● lower limit on m_1

● $\tilde{m}_0 = m_{up}$ \longrightarrow No LMA (SNO)
No $0, \beta\beta$

$\tilde{m}_0 = m_{down}$ \longrightarrow All solar solutions
No $0, \beta\beta$

$\tilde{m}_0 = m_{lepton}$ \longrightarrow All solar solutions
 $0, \beta\beta$ if $m_1 \gtrsim 10^3 eV$

● Future

● Inverse + Degenerate scheme

● δ, α, β and CP

● Analyze full LR leptogenesis
(Triplet-decay, χ, \dots)