

PATTERNS OF ν MASSES :

(a bottom-up approach) or

WHAT CAN WE LEARN ON/FROM U_{e3} ?

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Goals :

1. To reach a description as accurate as possible (cfr. hadronic sector) of
 - oscillation phenomena
 - $\bar{\nu} \nu$ processes (if measurable)
 - $\mu \rightarrow e \gamma$ in leptonic sector (if measurable)
2. Use the results in 1. to understand the origin of flavor in a fundamental theory.

Here: mainly 1. [2. \rightarrow Ramond, Lola, Yanagida, Altarelli]

- need of optimizing the experimental strategy
- need to define some basic theoretical framework.

3 main points enter the theoretical framework

① How many light states participate in the oscillations?

atm ν } do not require more than
solar ν } (ν_e, ν_μ, ν_τ)

→ check LSND (mini BOONE)

{ $\nu_e, \nu_\mu, \nu_\tau + \nu_s$
{ ν_e, ν_μ, ν_τ

↳ chosen here to proceed.

② Is $\Delta m_{solar}^2 \ll \Delta m_{atm}^2$ a good approximation?
most of the fits to atm use this

In this case:

atm $\leftrightarrow \Delta m_{atm}^2, \theta_{23}, \theta_{13}$
solar $\leftrightarrow \Delta m_{solar}^2, \theta_{12}, \theta_{13}$

| $\theta_{13} < 0.2$:
two almost decoupled systems

$\Delta m_{solar}^2 \ll \Delta m_{atm}^2$ supported by present data and essentially due to chlorine
without chlorine data Δm_{solar}^2 up to $10^{-3} eV^2$ is allowed (if $\sin^2 2\theta_{solar} \approx 1$)

→ SNO, Borexino, ...

③ $(\nu_e, \nu_\mu, \nu_\tau)$ are light Majorana fermions

mass term: $-\frac{1}{2} \nu^T m_\nu \nu$ $\nu = \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix}$
 $m_\nu^T = m_\nu$

From theory:

most convenient way to understand

$$|m_\nu| \ll |m_f| \quad f = u, d, l$$

$-\frac{1}{2} \nu^T m_\nu \nu$ arises from

after e.w. symmetry breaking: $\frac{(\varphi l)(\varphi l)}{\Lambda}$
 $\Lambda \rightarrow$ scale of Λ

$$m_\nu \approx \frac{v^2}{\Lambda}$$

From experiment:

no indication at present

- Oscillations cannot distinguish between pure Dirac and pure Majorana

- $0\nu\beta\beta$ occurs if ν are Majorana

present sensitivity: $|(m_\nu)_{ee}| < 0.2-0.8 \text{ eV}$

mini BOONE, SNO, Borexino, Kamland, ...

If ① + ② + ③, then

$$m_\nu = U m_{diag} U^T$$

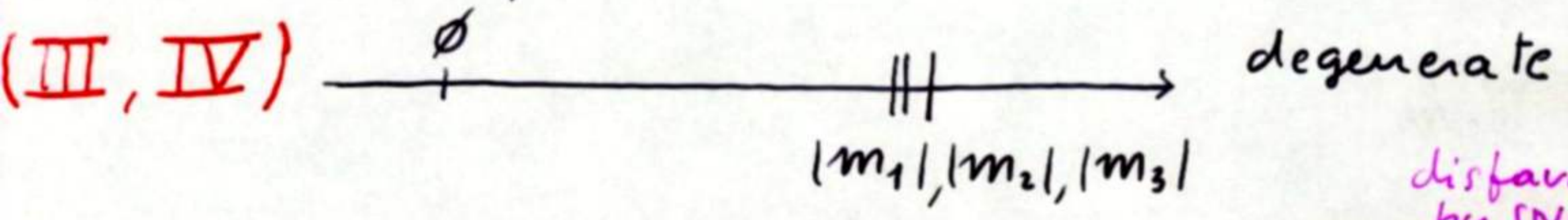
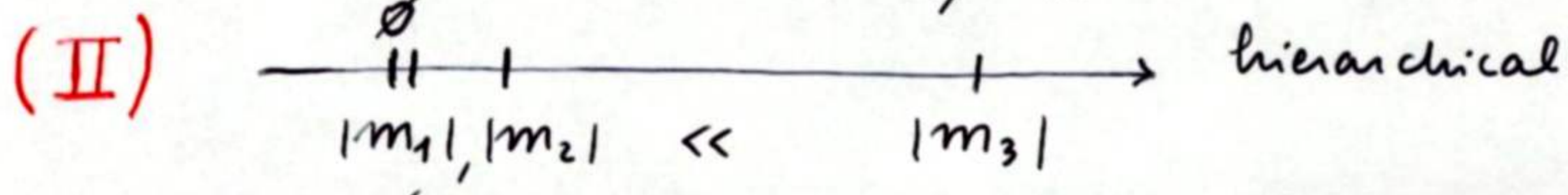
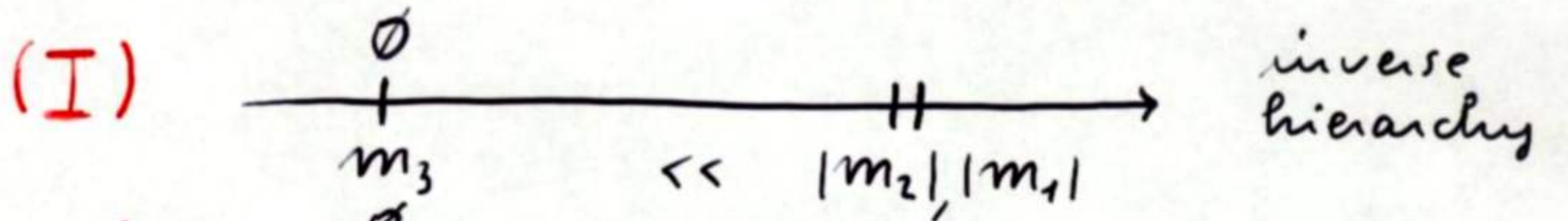
U ≡ mixing matrix in the lept. sector, if charged leptons are diagonal

④ ϕ^H order approximation:

- $\Delta m_{atm}^2 \approx 3 \times 10^{-3} eV^2$
 - $\sin^2 2\theta_{atm} \approx 1$
 - $\Delta m_{sol}^2 \approx \phi$
 - $U_{e3} \equiv \sin \theta_{13} = \phi$
 - $\delta \approx \phi$
- reasonable if ② holds

m_ν^o depends on:

(A) hierarchy among $|m_1|, |m_2|, |m_3|$:



disfavoured by SNO

- (B) $\begin{cases} \sin^2 2\theta_{12} \approx 0.0025 \approx \phi & \text{SAMSW} \\ \sin^2 2\theta_{12} \approx O(1) \approx 1 & \text{LAMS, LOW, VO} \end{cases}$

TABLE: \emptyset^{th} order TEXTURES

(up to phase redefinitions, in the basis where e, μ, τ are diagonal [exception IV])

Overall scale: m

m_{diag}	$\sin^2 2\theta_{12} \equiv 1$	$\sin^2 2\theta_{12} \equiv \emptyset$	Δm_{atm}^2
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(I)
 $[m, -m, \emptyset]$ $\frac{m}{\sqrt{2}} \begin{pmatrix} \emptyset & 1 & 1 \\ 1 & \emptyset & \emptyset \\ 1 & \emptyset & \emptyset \end{pmatrix}$ — m^2

(II)
 $[\emptyset, \emptyset, m]$ $\frac{m}{2} \begin{pmatrix} \emptyset & \emptyset & \emptyset \\ \emptyset & 1 & 1 \\ \emptyset & 1 & 1 \end{pmatrix} \leftrightarrow \text{SAME}$ m^2

(III)
 $[m, -m, m + \delta m]$
 $m \begin{pmatrix} \emptyset & 1/\sqrt{2} & 1/\sqrt{2} \\ 1/\sqrt{2} & \frac{1+\delta}{2} & -\frac{1+\delta}{2} \\ 1/\sqrt{2} & -\frac{1+\delta}{2} & \frac{1+\delta}{2} \end{pmatrix}$ — $2m^2\delta$

(IV)
 $[m, m, m + \delta m]$
 $m \begin{pmatrix} 1 & \emptyset & \emptyset \\ \emptyset & 1 & \emptyset \\ \emptyset & \emptyset & 1+\delta \end{pmatrix} \leftrightarrow \text{SAME}$ $2m^2\delta$

Other textures excluded because

- unstable under RGE

- not matching existing models

⑤ Up to now, $U_{e3} \equiv \emptyset$

Can we guess U_{e3} in (I) to (IV)?

Not without additional theoretical input:

- Regard (I) to (IV) as arising in some symmetry limit
- Add to (I), ... (IV) small SB terms to reproduce
 - $\begin{cases} \Delta m_{sol}^2 \neq \emptyset & (SA, LA, LOW, VO) \\ \sin^2 2\theta_{12} \neq \emptyset & (SA) \end{cases}$
- Look to what U_{e3} is generated.

Pay attention to:

5.1 $U \equiv U_l^\dagger U_\nu$ in real models

$\rightarrow U_{e3} \approx -\sin\theta_{12}^l \sin\theta_{atm} + \sin\theta_{13}^\nu$

can be either ν -dominated or l -dominated or both: $(\nu), (l), (\nu/l)$.

5.2 Results should be taken with great care:
They are indications, not theorems

Ultimately this exercise aims to answer the question:
What is the range of U_{e3} favored by current models?

TEXTURE (I):

- Barbieri, Hall, Smith, Strumia, Weiner '98

SYMMETRY LIMIT: $(L_e - L_\mu - L_\tau)$

FEATURES: $\theta_{23} \sim O(1)$ $\theta_{13} = \emptyset$ | bimaximal mixing
 $\theta_{12} = \pi/4$ (sharply)

AFTER SB: $m_\nu = \frac{m}{\sqrt{2}} \begin{pmatrix} \eta & 1 & 1 \\ 1 & \eta & \eta \\ 1 & \eta & \eta \end{pmatrix}$

$m_3 \approx m \eta$
 $m_1 \approx m(1 + c\eta)$
 $m_2 \approx -m(1 - c\eta)$

$\Delta m_{atm}^2 \approx m^2$
 $\Delta m_{sol}^2 \approx O(m^2 \eta)$

$\rightarrow \underbrace{\theta_{12}^\nu \approx \theta_{13}^l}_{(\nu/l)} \approx \eta \approx \frac{\Delta m_{sol}^2}{\Delta m_{atm}^2} = \begin{cases} 10^{-2} & \text{LA} \\ 10^{-4} - 10^{-5} & \text{LOW} \\ 10^{-7} & \text{VO} \end{cases}$

TEXTURE (II):

SYMMETRY LIMIT: - anomalous $U(1)$ B. Stech '98
- $U(2)$ - Barbieri, Gemmel, Romanino '99
- $SO(10)$ + HD operators - Sato, Yanagida '97
- R - Albright, Babu, Ban '98

FEATURES: $\theta_{23} \sim O(1)$
 $\theta_{13} = \emptyset$
 θ_{12} undetermined (2 \emptyset masses)
- Inger, Lavignac, Ramond '98
- Altarelli, F., Morina '99

well compatible with GUT ($SU(5)$, $SO(10)$) and see-saw. $\theta_{23} \sim O(1)$ may originate from l sector

large $(\mu-\tau)$ mixing \leftrightarrow large (s^c-b^c) mixing and then transferred to ν .

SYMMETRY BREAKING :

(II.1) SAMSW $\begin{cases} U_{e2} \approx \theta_{12}^\nu - \theta_{12}^l \cos \theta_{atm} \\ U_{e3} \approx \theta_{13}^\nu - \theta_{12}^l \sin \theta_{atm} \end{cases}$

(II.1.1) l -dominated $\begin{cases} U(2) \\ SO(10) + \text{HD operators} \end{cases}$

$U_{e3} \approx U_{e2} \tan \theta_{atm} \approx 10^{-2}$

(II.1.2) ν -dominated (often in models with $U(1)$)

$\rightarrow U_{e2} \approx \theta_{12}^\nu \quad U_{e3} \approx \theta_{13}^\nu$

$\frac{\theta_{13}^\nu}{\theta_{12}^\nu} \approx \frac{m_2}{m_3} \approx \sqrt{\frac{\Delta m_{sol}^2}{\Delta m_{atm}^2}}$

not a precise relation
order-of-magnitude estimate

$U_{e3} \approx \theta_{12}^\nu \sqrt{\frac{\Delta m_{sol}^2}{\Delta m_{atm}^2}} \approx 10^{-3} - 10^{-2}$

(II.2) $\theta_{12} \sim$ maximal LAMSW, LOW, VO

$\theta_{12} \approx \theta_{12}^\nu$ almost always

$\rightarrow \begin{cases} U_{e2} \approx \sin \theta_{12}^\nu - \theta_{12}^l \cos \theta_{12}^\nu \cos \theta_{atm} \\ U_{e3} \approx \theta_{13}^\nu - \theta_{12}^l \sin \theta_{atm} \end{cases}$

(II.2.1) l -dominated

$U_{e3} \approx -\theta_{12}^l \sin \theta_{atm} \ll \sqrt{\frac{m_e}{m_\mu}} \sin \theta_{atm} \approx 0.05$

(II.2.2) ν -dominated

$U_{e3} \approx \theta_{13}^\nu \approx \theta_{12}^\nu \sqrt{\frac{\Delta m_{sol}^2}{\Delta m_{atm}^2}} \approx \begin{cases} 10^{-1} & \text{LA} \\ 10^{-2} & \text{LOW} \\ 10^{-3} - 10^{-9} & \text{VO} \end{cases}$

TEXTURE (III)

SYMMETRY LIMIT: - SB $SO(3)$

FEATURES: $\theta_{23} \sim O(1)$
 $\theta_{13} = \emptyset$
 $\theta_{12} = \frac{\pi}{4}$ (sharply) | bimaximal mixing

- $m_{ee} = \emptyset$ limit from $0\nu\beta\beta$ evaded even for $m = (1 \div 2) eV$

- Stability under RGE depends on $(\frac{\epsilon}{\delta})$

$$\epsilon \approx \frac{y_\tau^2}{16\pi^2} \log \frac{\Lambda}{Q} \approx \left(\frac{10^{-6} \div 10^{-5}}{\cos^2 \beta} \right) \text{ in MSSM}$$

$$\delta \approx \frac{\Delta m_{atm}^2}{2m^2} \approx 1.5 \times 10^{-3} \left(\frac{1eV}{m} \right)^2$$

DEPARTURE FROM O^H APPROX:

Assume that is induced by RGE

$\theta_{12} = \pi/4$
 $\theta_{23} \sim O(1)$ } stable

$$\Delta m_{atm}^2 = 2m^2\delta$$

$$\Delta m_{sol}^2 = m^2\delta \left(\frac{\epsilon}{\delta} \right)^2$$

$$U_{e3} \approx \frac{\epsilon}{2\delta} = \frac{1}{\sqrt{2}} \sqrt{\frac{\Delta m_{sol}^2}{\Delta m_{atm}^2}} = \begin{cases} 10^{-1} \text{ LA} \\ 10^{-2} \text{ LOW} \\ 10^{-3} - 10^{-4} \text{ VO} \end{cases}$$

$$\Delta m_{sol}^2 \approx 600 \left(\frac{m}{1eV} \right)^4 \epsilon^2 eV^2$$

$\rightarrow \epsilon \approx \begin{cases} \text{LA} & (10^{-3} \div 10^{-4}) \left(\frac{1eV}{m} \right)^2 & o.k. \\ \text{LOW} & (10^{-4} - 10^{-5}) \left(\frac{1eV}{m} \right)^2 & \rightarrow \tan \beta < 10 \\ \text{VO} & 10^{-6} \left(\frac{1eV}{m} \right)^2 & \rightarrow m \approx 0.1eV \end{cases}$

TEXTURE (IV)

SYMMETRY LIMIT: { FLAVOR DEMOCRACY
S_{3L} x S_{3R} , O(3)_L x O(3)_R

FEATURES :

charged sector: $m_f = \frac{m_f^0}{3} \begin{pmatrix} 1 & 1 & 1 \\ 1 & 1 & 1 \\ 1 & 1 & 1 \end{pmatrix} + \delta m_f$
(f = u, d, l)

→ hierarchy of charged fermion masses
(1 dominant mass x family)

→ small angles in V_{CKM} from cancellation
in $U_u^+ U_d$.

ν-sector $m_\nu = m_\nu^0 \begin{pmatrix} 1 & \emptyset & \emptyset \\ \emptyset & 1 & \emptyset \\ \emptyset & \emptyset & 1 \end{pmatrix} + \delta m_\nu$

→ large angles in $U = U_e^+ U_\nu$ $\sin^2 2\theta_{atm} = \frac{8}{9}!$

attention to:

- $(m_{\nu ee} \sim m_\nu^0 \lesssim (0.2-0.8) eV$ from $\emptyset \nu \beta \beta$
- RGE tends to destabilize the texture.

SYMMETRY BREAKING:

on a empirical basis:

$$\delta m_\nu = \begin{pmatrix} 0 & \eta & 0 \\ \eta & 0 & 0 \\ 0 & 0 & \delta \end{pmatrix} \rightarrow \text{SAMSW}$$

$$|\eta| \ll |\delta| \ll m_\nu^0$$

$$\delta m_\nu = \begin{pmatrix} 0 & 0 & 0 \\ 0 & \eta & 0 \\ 0 & 0 & \delta \end{pmatrix} \rightarrow \begin{cases} \text{LA} \\ \text{LOW} \\ \text{VO} \end{cases}$$

In any case U_{e3} is l-dominated:

$$U_{e3} \approx \underbrace{g_{12}^l}_{\sqrt{2/3}} \underbrace{\sin \theta_{atm}}_{\sqrt{2/3}} \equiv \sqrt{\frac{2}{3}} \sqrt{\frac{m_e}{m_\mu}} \approx 0.05$$

spoiled by running

TIMETABLE (very rough...)

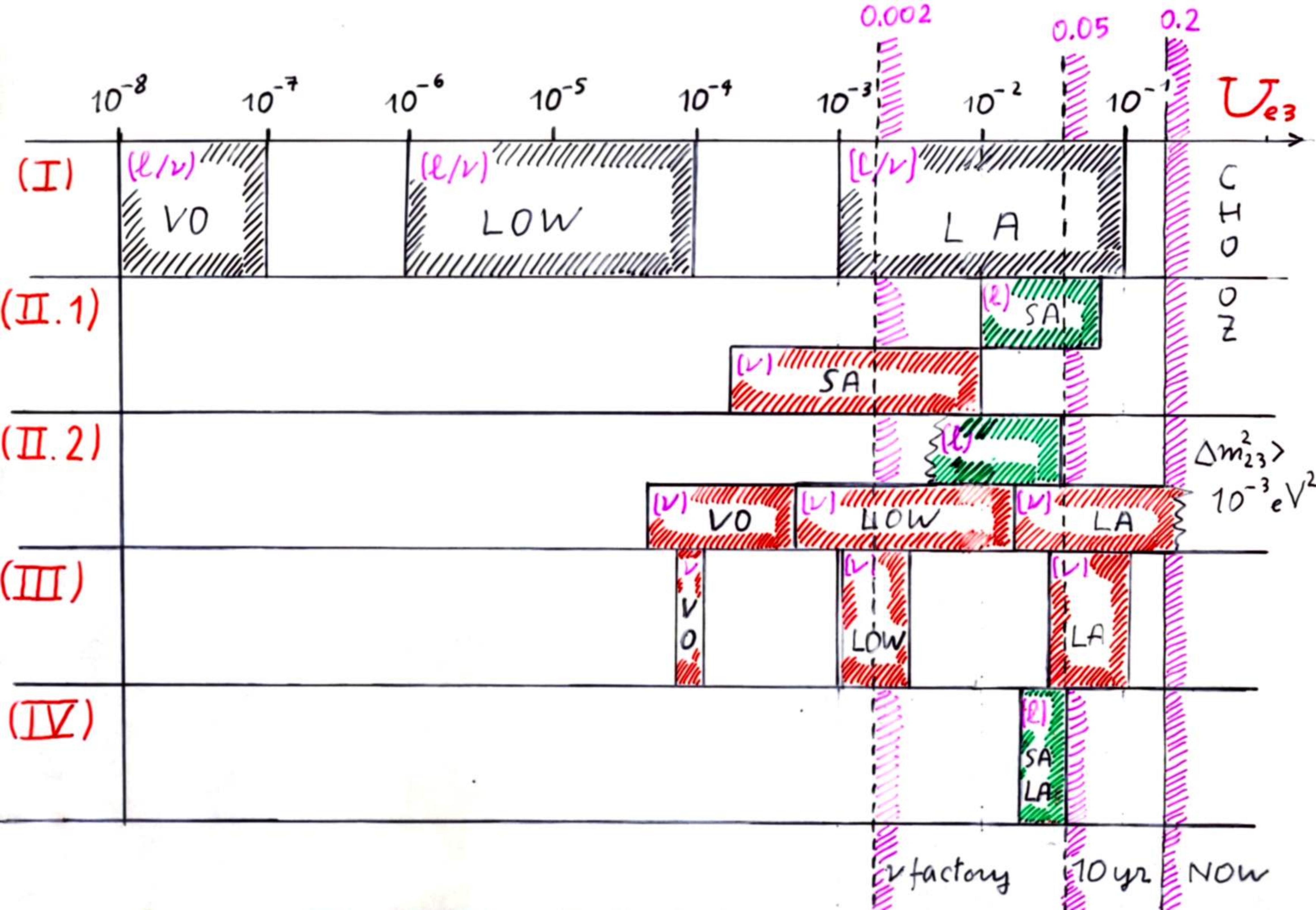
		today	in 10 yr	ν factory
<u>atm</u>	Δm_{23}^2 $\sin^2 2\theta_{23}$	$\sim 10^{-3} eV^2$ ~ 1	10% accuracy K2K, MINOS, NGS	5% accuracy
<u>solar</u>	Δm_{12}^2 $\sin^2 2\theta_{12}$	$\leq 10^{-4} eV^2$?	10-20% accuracy SNO, Borexino, Kamland	
	$\sin^2 2\theta_{13}$ δ_{CP} Δm_{23}	< 0.1 ? ?	< 0.01 JHF. ? ?	< 0.00002 Yes if LA accessible
	$ U_{e3} $	< 0.15	< 0.05	< 0.002

To plot U_{e3} I used:

Δm_{solar}^2	$4 \times 10^{-6} - 10^{-5}$	$10^{-5} - 10^{-4}$	$10^{-8} - 10^{-7}$	10^{-10}
$\sin^2 2\theta_{solar}$	$10^{-3} - 10^{-2}$	0.6-0.9	0.6-0.9	0.6-0.9
	SAMSW	LAMSW	LOW	VO

When the theoretical estimate suggests:

$\approx \rightarrow$ factor 10 of uncertainty $\equiv \left(\frac{3 \times \text{upper}}{\frac{1}{3} \times \text{lower}} \right)$



COMMENTS

1. With the planned experiments, we will have access to a significant portion of the θ_{e3} parameter space.

2. A knowledge of the absolute spectrum would not help in clarifying θ_{e3} and viceversa the value of θ_{e3} by itself would not help in discriminate among (I) - (IV)

3. solar ν and θ_{e3}

$\theta_{e3} \leftarrow \dots \rightarrow$ solution to solar ν

- Values of θ_{e3} close to 0.1 seem possible only within **LA** [I, II.2, III]

- $0.02 \leq \theta_{e3} \leq 0.1 \rightarrow$ **LA** still preferred but also **SA** [II.1, IV] possible

Exception: hierarchical, $\sin^2 2\theta_{12} \sim 1$ if l -dominated

- **VO** seems related to an unmeasurable θ_{e3} . Exception: as before

- $0.001 \leq \theta_{e3} \leq 0.02$ difficult region

SA, LA & LOW possible with ν -dominance preferred